
Stellar Astronomy and Astrophysics (SS12)

Stefan Jordan and Ralf Klessen

Exercise 5 for May 29, 2012

1. Convection:

In the lecture the Schwarzschild criterion for convection was derived. Convection occurs, if the logarithmic temperature gradient $\nabla = \left(\frac{d \ln T}{d \ln P}\right)$ for radiation is larger than that for the adiabatic one:

$$\nabla_{\text{rad}} > \nabla_{\text{ad}} = 1 - \frac{1}{\gamma} \quad (1)$$

with $\gamma = \frac{c_p}{c_v}$ being the ratio between the specific heats. This equation was derived under the assumption that $\frac{d\rho_{\text{blob}}}{dr} < \frac{d\rho_{\text{sur}}}{dr}$ is equivalent to $\frac{dT_{\text{blob}}}{dr} > \frac{dT_{\text{sur}}}{dr}$. For this argument we used the equation of state, which in the case of an ideal gas is $P = \frac{\rho R}{\mu} T$. However, the argument is only valid if the mean molecular weight μ is constant.

Assume now that due to nuclear burning there is a chemical gradient in the star (μ is a function of the radial coordinate r). How would the convection criterion change if we assume that such a chemical gradient exists?

Ansatz: $\frac{d\rho}{\rho} = \lim_{dr \rightarrow 0} \frac{\rho(r+dr) - \rho(r)}{\rho(r)}$. Convection occurs, if the density $\rho_{\text{blob}}(r+dr) = \rho_{\text{ad}}(r+dr)$ is smaller than $\rho_{\text{sur}}(r+dr)$. Further assume that the chemical composition in the blob remains constant $\mu_{\text{blob}}(r+dr) = \mu_{\text{blob}}(r)$, while it is changing in the surrounding $\mu_{\text{sur}}(r) \neq \mu_{\text{sur}}(r+dr)$.

Is convection easier or suppressed if $\mu_{\text{sur}}(r+dr) < \mu_{\text{sur}}(r)$? Do you see any problem using the modified convection criterion?

2. Hydrostatic Equation:

Show that the equation of hydrostatic equilibrium may be written as

$$\frac{dP}{d\tau} = \frac{g}{\kappa} \quad (2)$$

with P being the pressure, g being the local gravitational acceleration and τ being the optical depth defined for the mass absorption coefficient κ .

3. Pressure in a white dwarf:

- Estimate the density of a white dwarf that has a solar mass packed into a sphere with about 0.01 solar radii.
- Estimate the degeneracy pressure in the interior of this white dwarf and compare this pressure with the thermal pressure of a gas that has a temperature of 10^7 K.