# Assignment \#1: due Tuesday, Oct. 23 

## Theoretical Astrophysics

Winter 2006/2007
lecturer: Prof. Ralf Klessen, ITA Albert-Ueberle-Str. 2

## 1. Number of stars in Milky Way

Does the Milky Way Galaxy contain more stars than there are grains of sand in the beach volleyball court at the Neckarwiese? Please justify your answer using simple order-of-magnitude estimates.

## 2. Timescale estimates for the Sun

(a) Dynamical timescale: The collapse timescale for a self-gravitating object is given by $t_{\mathrm{dyn}} \approx 1 / \sqrt{G \rho}$. Calculate it for the Sun assuming a mean density of $\rho=$ $1.4 \mathrm{~g} \mathrm{~cm}^{-3}$
(b) Sound crossing time: The sound speed in the solar core is roughly $350 \mathrm{~km} \mathrm{sec}^{-1}$. How long does a sound wave need to cross the Sun assuming a constant sound speed throughout the Sun (the solar radius is $6.96 \times 10^{5} \mathrm{~km}$ ).
(c) Nuclear timescale: The Sun's energy is produced by the process of fusion of hydrogen into helium. If $10 \%$ of the solar mass is consumed in this process during the Sun's lifetime, how long does the Sun's energy production persist if the Sun's energy loss (i.e. luminosity: $L_{\odot}=3.846 \times 10^{33} \mathrm{ergs} \mathrm{sec}^{-1}$ ) is constant during that time. Use the formula $\tau=E / \dot{E}$ to estimate the nuclear time scale. Note that $0.7 \%$ of the hydrogen rest mass is turned into energy in the fusion process.
(d) Kelvin-Helmholtz timescale: The Kelvin-Helmholtz timescale is the ratio of the gravitational energy of an object to its luminosity. Calculate the Kelvin-Helmholtz timescale for the Sun, assume the Sun has a constant density.
3. Angular momentum
(a) Estimate the total angular momentum of a $0.1 \mathrm{pc}\left(\mathrm{pc}=3.085 \times 10^{18} \mathrm{~cm}\right)$ size molecular cloud core with a mean density of $\rho=1.67 \times 10^{-20} \mathrm{~g} \mathrm{~cm}^{-3}$ and a constant angular velocity of $\Omega=10^{-14} \mathrm{rad} \mathrm{sec}^{-1}$.
(b) Assuming the above cloud collapses to a single solar type star, what would be the rotational velocity at its surface (solar mass $M_{\odot}=1.989 \times 10^{33} \mathrm{~g}$ )? Could gravity hold this object together? Discuss this result.
(c) Calculate the total angular momentum of the Sun assuming a mean rotational period of 30 days.

