## Assignment #7: due Tuesday, Dec. 4

## **Theoretical Astrophysics**

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## 1. Parker instability

Consider an isothermal gas in the galactic disk which is threaded with a horizontal magnetic field. Assume a constant gravitational field perpendicular to the disk plane in the z direction, i.e.  $\vec{g} = -\hat{z}g$  and a magnetic field parallel to the disk plane x which varies only with z, i.e.  $B = \hat{x}B(z)$ . For simplicity study the system in two dimensions using cartesian coordinates.

(a) Assume that the system is in hydrostatic equilibrium with a constant ratio of the magnetic to thermal pressure, i.e.

$$\alpha \equiv \frac{B^2}{8\pi P} = \text{const.} \tag{1}$$

30 pt

What is the pressure distribution as a function of z? Use the relation  $P = c_s^2 \rho$  where c is the constant speed of sound and the scale height  $H = (1 + \alpha) c^2/g$  to express the result.

Now consider this system slightly perturbed out of its equilibrium. Then, from the linear perturbation analysis one gets the following dispersion relation in the xz-plane,

$$n^{4} + c_{\rm s}^{2} \left[ (1+2\alpha) \left( k^{2} + \frac{k_{0}^{2}}{4} \right) \right] n^{2} + k_{x}^{2} c_{\rm s}^{4} \left[ 2\alpha k^{2} + k_{0}^{2} \left[ \left( 1 + \frac{3\alpha}{2} \right) - (1+\alpha)^{2} \right] \right] = 0 , \qquad (2)$$

(you might want to derive this relation in your spare time), where  $n = i\omega$ ,  $k_0 = H^{-1}$ , and the Fourier modes in the x and z direction for the perturbed quantities are

$$\exp(i\omega t + ik_x x) \quad , \quad \exp(i\omega t + ik_z z) \tag{3}$$

with  $k^2 = k_x^2 + k_z^2$ .

(b) Show that in the absence of a magnetic field all roots (in terms of  $n^2$ ) of this dispersion relation are negative, i.e.  $n^2 < 0$ . What is the physical implication of this result regarding the instability?

(c) In the case of a non-vanishing magnetic field derive the instability criterion for the Parker instability (magnetic Rayleigh-Taylor instability)

$$\left(\frac{k}{k_0/2}\right)^2 < 2\alpha + 1 . \tag{4}$$

*Hint:* Use the roots of  $n^2$  to find at leat one unstable mode, i.e.  $n^2 < 0$ .

(d) Show that the instability criterion is equivalent to

$$\lambda_x > \Lambda_x \equiv 4\pi H \left[\frac{1}{2\alpha + 1}\right]^{1/2} \tag{5}$$

$$\lambda_z > \Lambda_z \equiv \frac{\Lambda_x}{\left(1 - \left(\Lambda_x/\lambda_x\right)^2\right)^{1/2}}, \qquad (6)$$

with the wavelengths  $\lambda_x = 2\pi/k_x$ , and  $\lambda_z = 2\pi/k_z$ .