# Assignment \#7: due Tuesday, Dec. 4 

## Theoretical Astrophysics

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## 1. Parker instability

30 pt

Consider an isothermal gas in the galactic disk which is threaded with a horizontal magnetic field. Assume a constant gravitational field perpendicular to the disk plane in the $z$ direction, i.e. $\vec{g}=-\hat{z} g$ and a magnetic field parallel to the disk plane $x$ which varies only with $z$, i.e. $B=\hat{x} B(z)$. For simplicity study the system in two dimensions using cartesian coordinates.
(a) Assume that the system is in hydrostatic equilibrium with a constant ratio of the magnetic to thermal pressure, i.e.

$$
\begin{equation*}
\alpha \equiv \frac{B^{2}}{8 \pi P}=\text { const. } \tag{1}
\end{equation*}
$$

What is the pressure distribution as a function of $z$ ? Use the relation $P=c_{\mathrm{s}}^{2} \rho$ where $c$ is the constant speed of sound and the scale height $H=(1+\alpha) c^{2} / g$ to express the result.

Now consider this system slightly perturbed out of its equilibrium. Then, from the linear perturbation analysis one gets the following dispersion relation in the $x z$-plane,

$$
\begin{align*}
n^{4}+ & c_{\mathrm{s}}^{2}\left[(1+2 \alpha)\left(k^{2}+\frac{k_{0}^{2}}{4}\right)\right] n^{2}+ \\
& k_{x}^{2} c_{\mathrm{s}}^{4}\left[2 \alpha k^{2}+k_{0}^{2}\left[\left(1+\frac{3 \alpha}{2}\right)-(1+\alpha)^{2}\right]\right]=0 \tag{2}
\end{align*}
$$

(you might want to derive this relation in your spare time), where $n=i \omega, k_{0}=$ $H^{-1}$, and the Fourier modes in the $x$ and $z$ direction for the perturbed quantities are

$$
\begin{equation*}
\exp \left(i \omega t+i k_{x} x\right) \quad, \quad \exp \left(i \omega t+i k_{z} z\right) \tag{3}
\end{equation*}
$$

with $k^{2}=k_{x}^{2}+k_{z}^{2}$.
(b) Show that in the absence of a magnetic field all roots (in terms of $n^{2}$ ) of this dispersion relation are negative, i.e. $n^{2}<0$. What is the physical implication of this result regarding the instability?
(c) In the case of a non-vanishing magnetic field derive the instability criterion for the Parker instability (magnetic Rayleigh-Taylor instability)

$$
\begin{equation*}
\left(\frac{k}{k_{0} / 2}\right)^{2}<2 \alpha+1 \tag{4}
\end{equation*}
$$

Hint: Use the roots of $n^{2}$ to find at leat one unstable mode, i.e. $n^{2}<0$.
(d) Show that the instability criterion is equivalent to

$$
\begin{align*}
& \lambda_{x}>\Lambda_{x} \equiv 4 \pi H\left[\frac{1}{2 \alpha+1}\right]^{1 / 2}  \tag{5}\\
& \lambda_{z}>\Lambda_{z} \equiv \frac{\Lambda_{x}}{\left(1-\left(\Lambda_{x} / \lambda_{x}\right)^{2}\right)^{1 / 2}} \tag{6}
\end{align*}
$$

with the wavelengths $\lambda_{x}=2 \pi / k_{x}$, and $\lambda_{z}=2 \pi / k_{z}$.

