

Assignment #8: due Tuesday, Dec. 15

Theoretical Astrophysics

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Ralf Klessen & Ingo Berentzen, ZAH/ITA, Albert-Ueberle-Str. 2, 69120 Heidelberg

In a cold, magnetized plasma consisting of electrons (charge $q_e = -e$, mass m_e) and ions (charge $q_i = Ze$, mass m_i), the equation governing the propagation of a wave-like disturbance, $\vec{E} = \vec{E}_0 \exp(i\vec{k} \cdot \vec{x} - i\omega t)$ is

$$\mathcal{E} \vec{E} = 0. \quad (1)$$

We use cartesian coordinates with basis $(\vec{e}_x, \vec{e}_y, \vec{e}_z)$ and assume that the wave propagates along the magnetic field which we take parallel to \vec{e}_z . In this case, the matrix \mathcal{E} is

$$\mathcal{E} = \begin{pmatrix} S - n^2 & -iD & 0 \\ iD & S - n^2 & 0 \\ 0 & 0 & P \end{pmatrix}, \quad (2)$$

where $n = kc/\omega$ is the refractive index, and

$$S = 1 - \frac{\omega_{pe}^2}{\omega^2 - \Omega_e^2} - \frac{\omega_{pi}^2}{\omega^2 - \Omega_i^2}, \quad (3)$$

$$D = \frac{\omega_{pe}^2 \Omega_e}{\omega(\omega^2 - \Omega_e^2)} + \frac{\omega_{pi}^2 \Omega_i}{\omega(\omega^2 - \Omega_i^2)}, \quad (4)$$

$$P = 1 - \frac{\omega_{pe}^2}{\omega^2} - \frac{\omega_{pi}^2}{\omega^2}. \quad (5)$$

The quantities $\omega_{pe,pi} = \sqrt{4\pi n_{e,i} q_{e,i}^2 / m_{e,i}}$ are the electron and ion plasma frequencies (with $n_{e,i}$ the number densities) and $\Omega_{e,i} = q_{e,i} B / m_{e,i} c$ are the electron and ion gyration frequencies. Note that both have opposite signs.

1. Alfvén waves

30 pt

- Find the dispersion relation in a neutral electron-proton plasma in the low frequency limit, $\omega \ll \Omega_i$ and $\omega \ll \omega_{pi}$. Make use of the fact that $m_i \gg m_e$. Show that only transversal waves are permitted.
- Find the polarization vectors of the corresponding transversal modes. Note, they correspond to the eigenvectors of the system.

2. Faraday rotation

30 (+ 5) pt

- (a) In the high frequency limit, $\omega \gg \omega_{pe,pi}$ and $\omega \gg \Omega_{e,i}$, show that the dispersion relation in the electron-proton plasma for waves travelling in the positive z direction can be written approximately as

$$\frac{kc}{\omega} = 1 - \frac{\omega_{pe}^2 + \omega_{pi}^2}{2\omega^2} \pm \frac{\omega_{pe}^2 \Omega_e}{2\omega^3}. \quad (6)$$

The upper and lower signs refer to the polarization vectors $(1/\sqrt{2}, \pm i/\sqrt{2}, 0)$. Use again the fact that $m_i \gg m_e$.

- (b) Show that a linearly polarized photon that is emitted along the magnetic field will rotate its direction of polarization as it propagates by an amount proportional to the inverse square of its frequency.
- (c) For ionized hydrogen gas in the Galactic plane with $n = 1 \text{ cm}^{-3}$ and $B = 20 \mu\text{G}$, find the distance over which a photon of frequency 3 GHz that is emitted linearly polarized in the x -direction travels before it is converted to one polarized in the y -direction. Assume propagation along a uniform magnetic field.
- (d) How does the result change if the photon propagates in a hypothetical electron-positron plasma, where $m_e = m_i$? (+5 bonus points)