

Assignment #3: due Thursday, Nov. 4, 2010

Theoretical Astrophysics

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Ralf Klessen, ZAH/ITA, Albert-Ueberle-Str. 2, 69120 Heidelberg

1. Equation of hydrostatic balance (bonus points 10 pt)

Using the equation of hydrostatic balance, obtain a crude estimate of the central temperature of the Sun. Hint: Approximate the differential operator by the finite difference between the solar surface and center.

2. Degenerate electron gas 25 pt

Consider a degenerate electron gas at zero temperature. That means that all quantum levels up to the Fermi momentum p_F are filled and all others are empty. The phase-space probability density is then given by

$$f(\vec{q}, \vec{p}) = \begin{cases} 2/h^3 & \text{for } p \leq p_F \\ 0 & \text{for } p > p_F \end{cases} \quad (1)$$

The factor 2 takes into account the two spin orientations of the electrons.

Use this distribution function to compute the density and pressure and show that this gas has a polytropic equation of state

$$P = K\rho^\gamma \quad (2)$$

Find the index γ and the constant K for the non-relativistic case. Assume that each electron is accompanied by one proton whose kinetic energy is negligible.

3. Lane-Emden equation 30 pt

In a spherically symmetric system, the equations of hydrostatic equilibrium and Poisson's equation are:

$$\frac{1}{\rho} \frac{dP}{dr} = -\frac{d\Phi}{dr} \quad (3)$$

$$\frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{d\Phi}{dr} \right) = 4\pi G \rho \quad (4)$$

where Φ is the gravitational potential and G Newton's gravitational constant.

- (a) Taking $\Phi(r_{\text{surf}}) = 0$ and $\rho(r_{\text{surf}}) = 0$ at the surface of the star, $r = r_{\text{surf}}$, show that for a polytropic equation of state, i.e. $P = K\rho^{(n+1)/n} = K\rho^\gamma$, the density in the star ($\Phi < 0$) can be expressed as

$$\rho = \left(\frac{-\Phi}{(n+1)K} \right)^n \quad (5)$$

- (b) Substitute this expression into the Poisson equation and show that this then reduces to the *Lane-Emden equation*:

$$\frac{1}{z^2} \frac{d}{dz} \left(z^2 \frac{dw}{dz} \right) + w^n = 0 \quad (6)$$

when written in terms of the variables $w = \Phi/\Phi_c$ and $z = r/r_0$, where Φ_c is the potential at $r = 0$ and r_0 is a characteristic length scale. Find an expression for r_0 in terms of n and K .

- (c) Given that there exists a solution of the Lane-Emden equation for the given system, show that the radius R of a non-relativistic degenerate star ($n = 3/2$) is related to its total mass by

$$R \propto M^{-1/3}. \quad (7)$$