HII Regions from "Exiled" Massive Stars

J. Mackey¹, V.V. Gvaramadze^{2,3}, N. Langer¹, D. M.-A. Meyer¹

- 1. Argelander-Institut für Astronomie, Universität Bonn.
- 2. Sternberg Astronomical Institute, Moscow; 3. Isaac Newton Institute of Chile, Moscow Branch.
 - 15-25% of O stars are not in clusters, and (almost) all of these are "exiles", moving through the ISM away from their birthplace.
 - Their winds and HII regions can have strong dynamical effects on the ISM.
 - We study HII regions from these stars with 2D and 3D (magneto)hydrodynamic simulations.
 - Kinetic energy input to ISM from HII regions can be comparable to that from winds.
 - Ionisation front can be dynamically unstable.
 - Modelling can be used to constrain mass-loss rates of O stars.