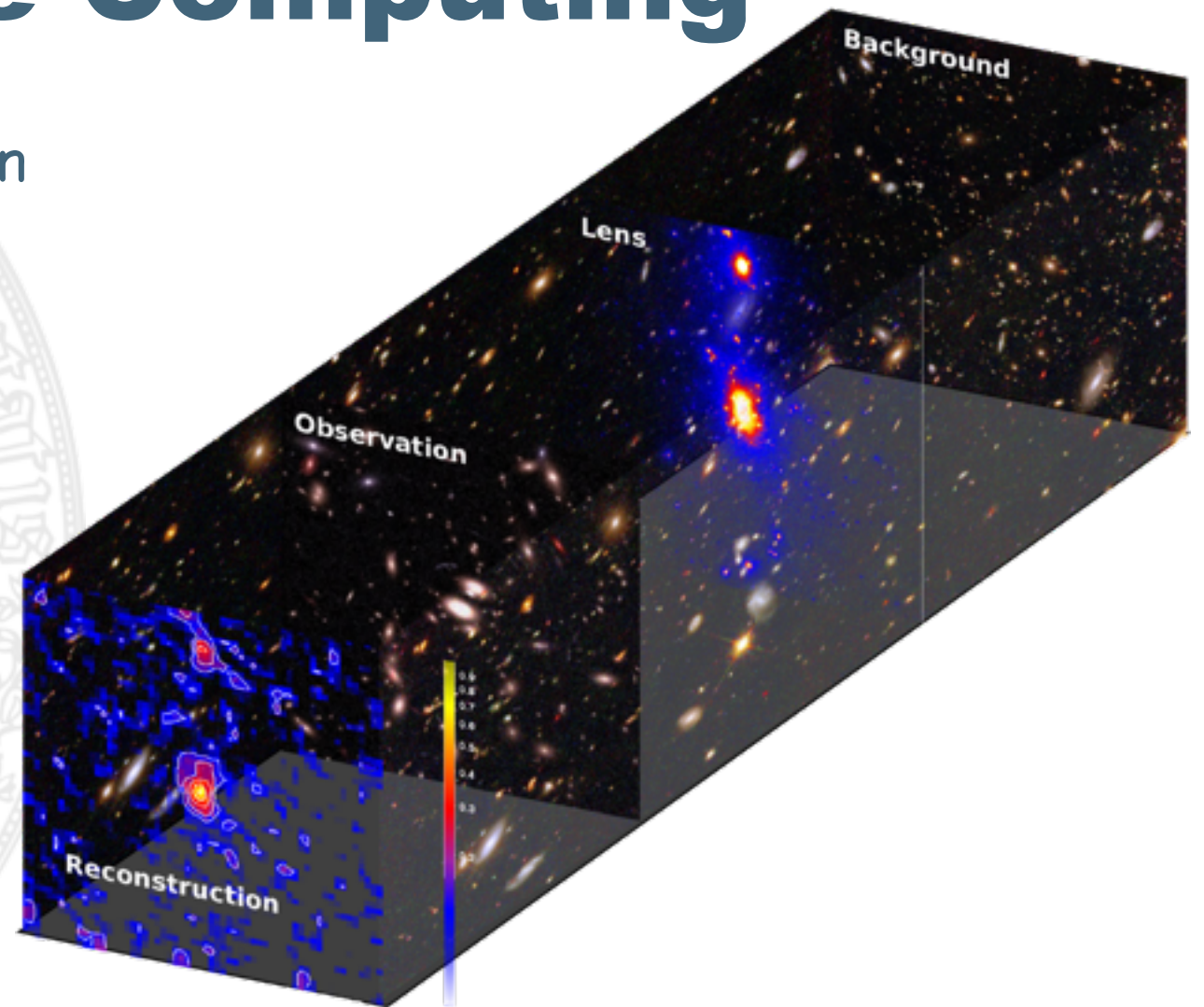


Cosmology with Galaxy Clusters: Observations meet High-Performance-Computing

Julian Merten
(ITA/ZAH)

- **Clusters of galaxies**
- **GPU lensing codes**
- **Abell 2744**
- **CLASH: A HST/MCT programme**



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Clusters of galaxies

- ▶ DM dominated

$\sim 10^{15} M_{\odot}$, Mpc scale

- ▶ Baryonic component not dominant, though not negligible
- ▶ Observable in 3 regimes
- ▶ Cosmic laboratories
- ▶ Strong grav. lenses
- ▶ Still pose riddles



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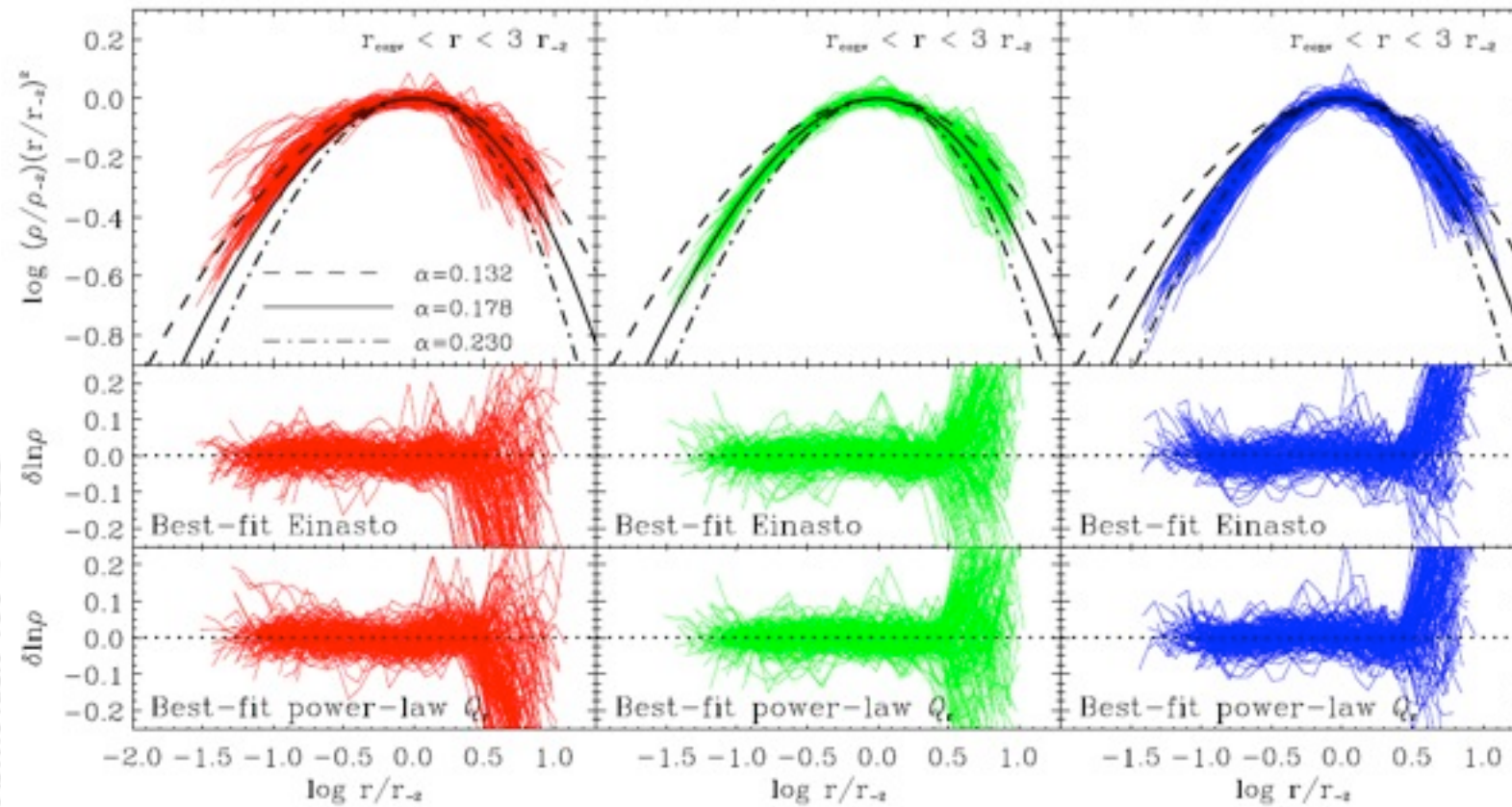


Cluster puzzles



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Cluster puzzles

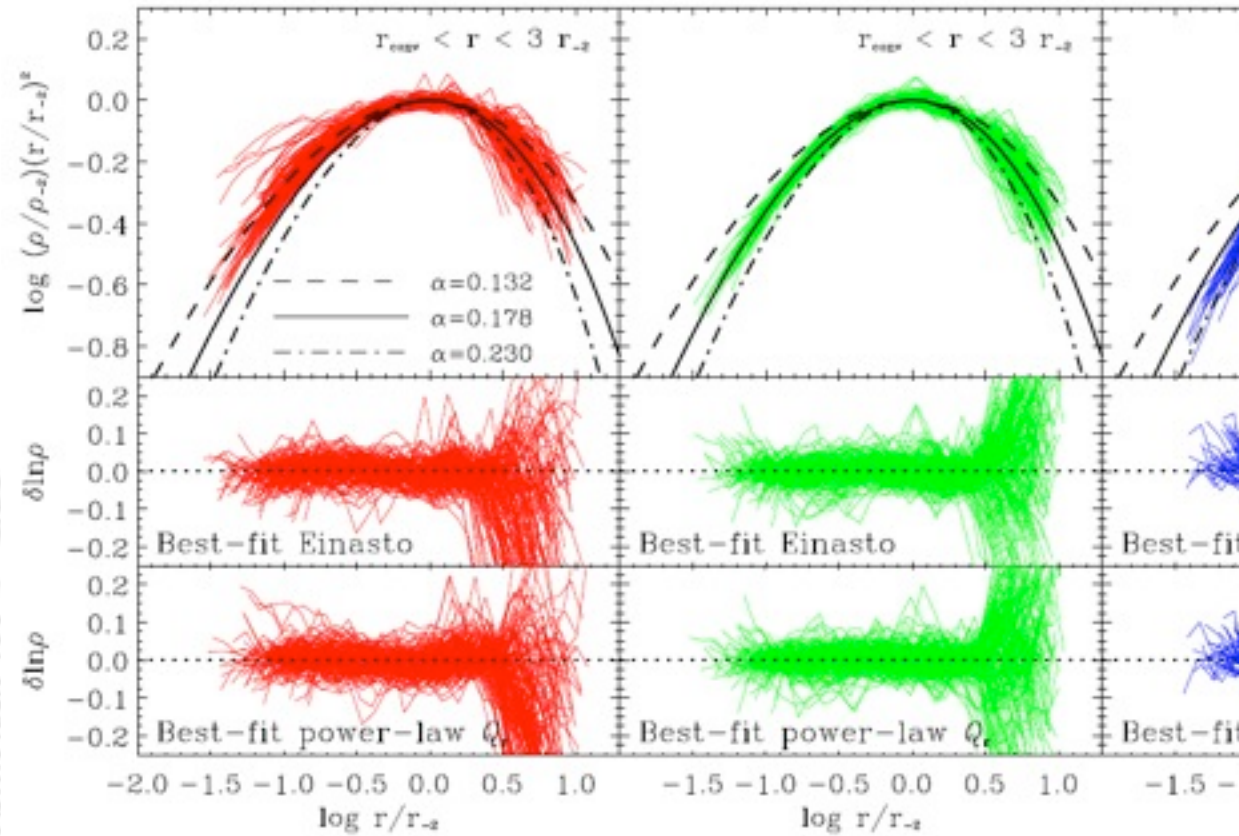


Ludlow11+

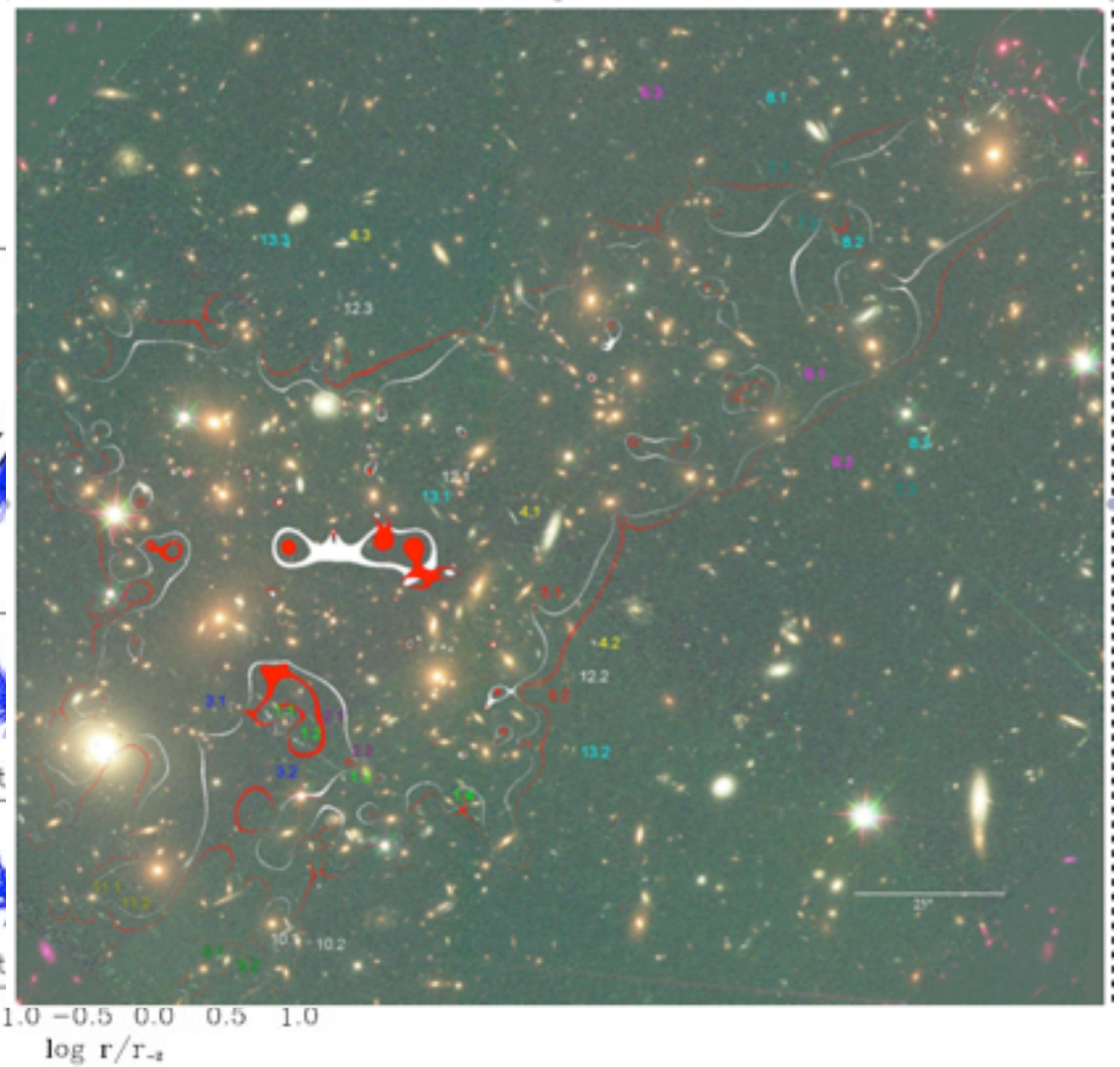
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Cluster puzzles



Ludlow11+

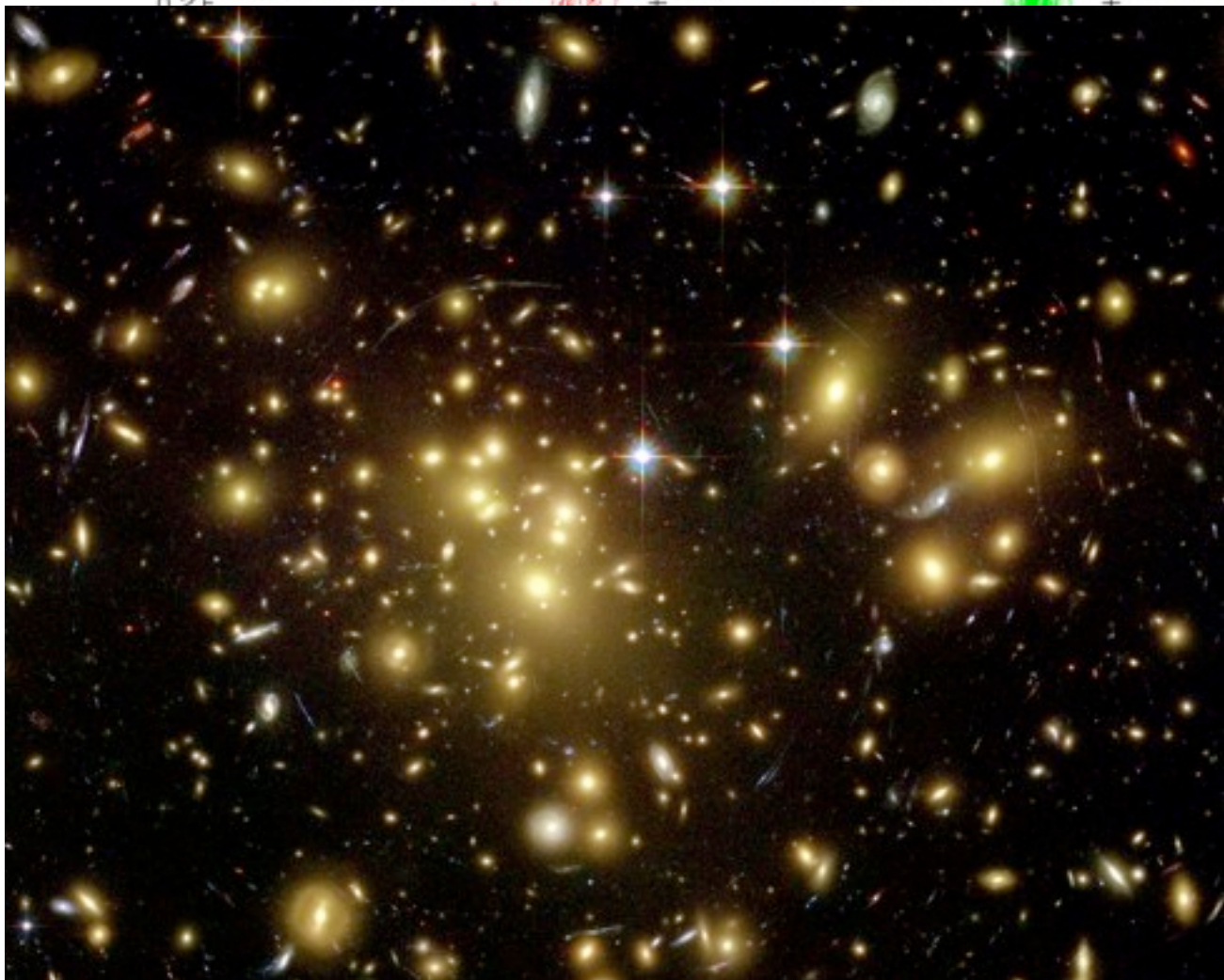
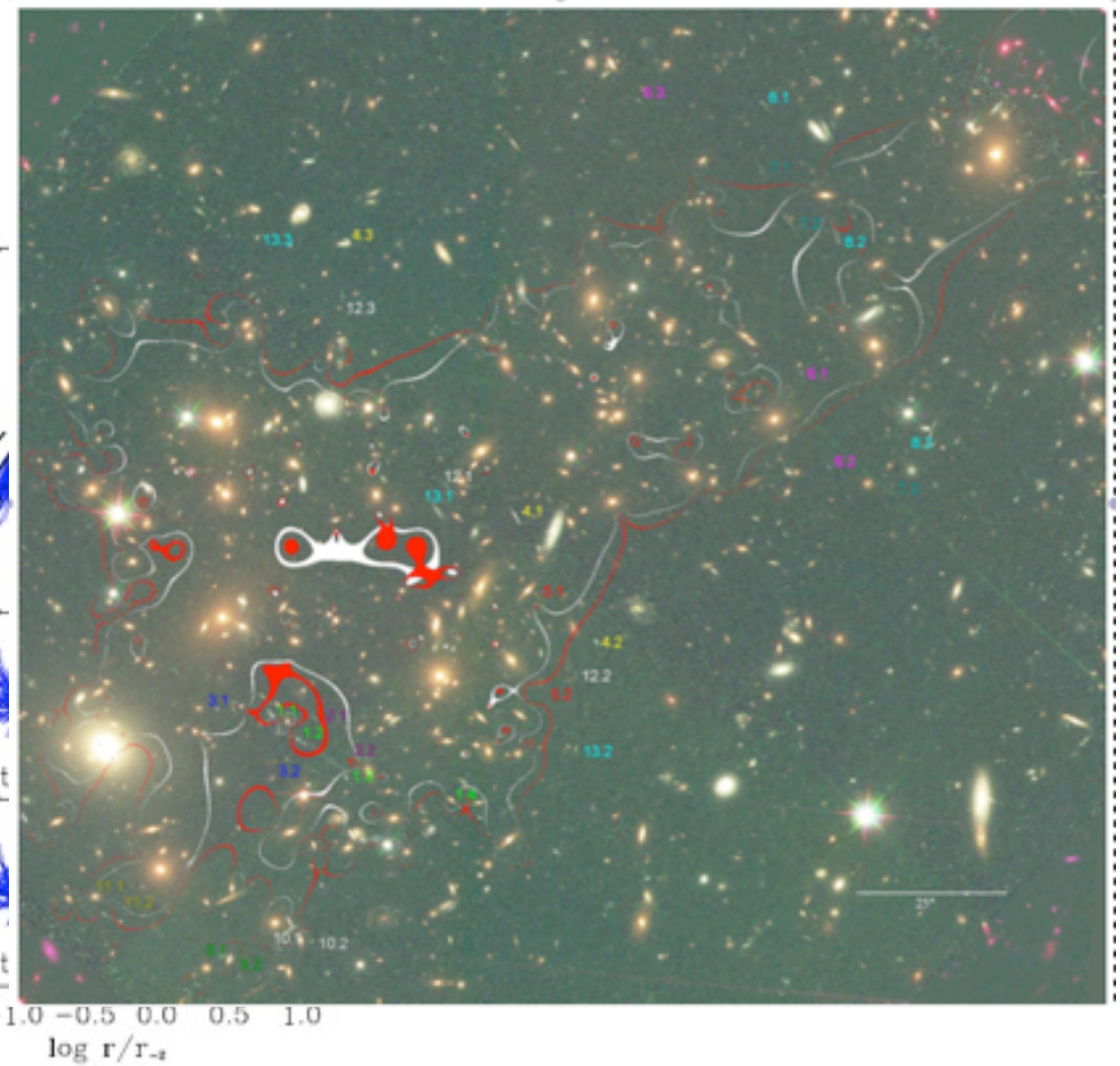
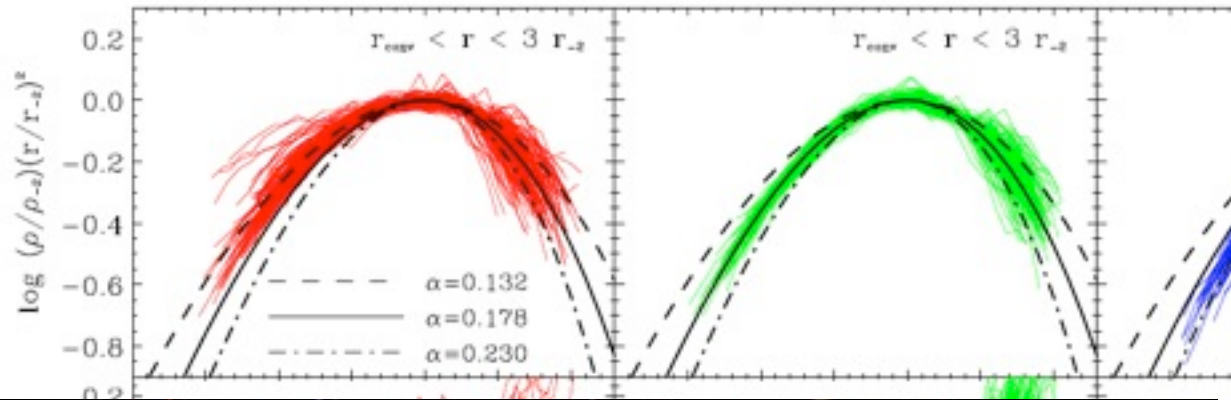


Zitrin10+

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Cluster puzzles



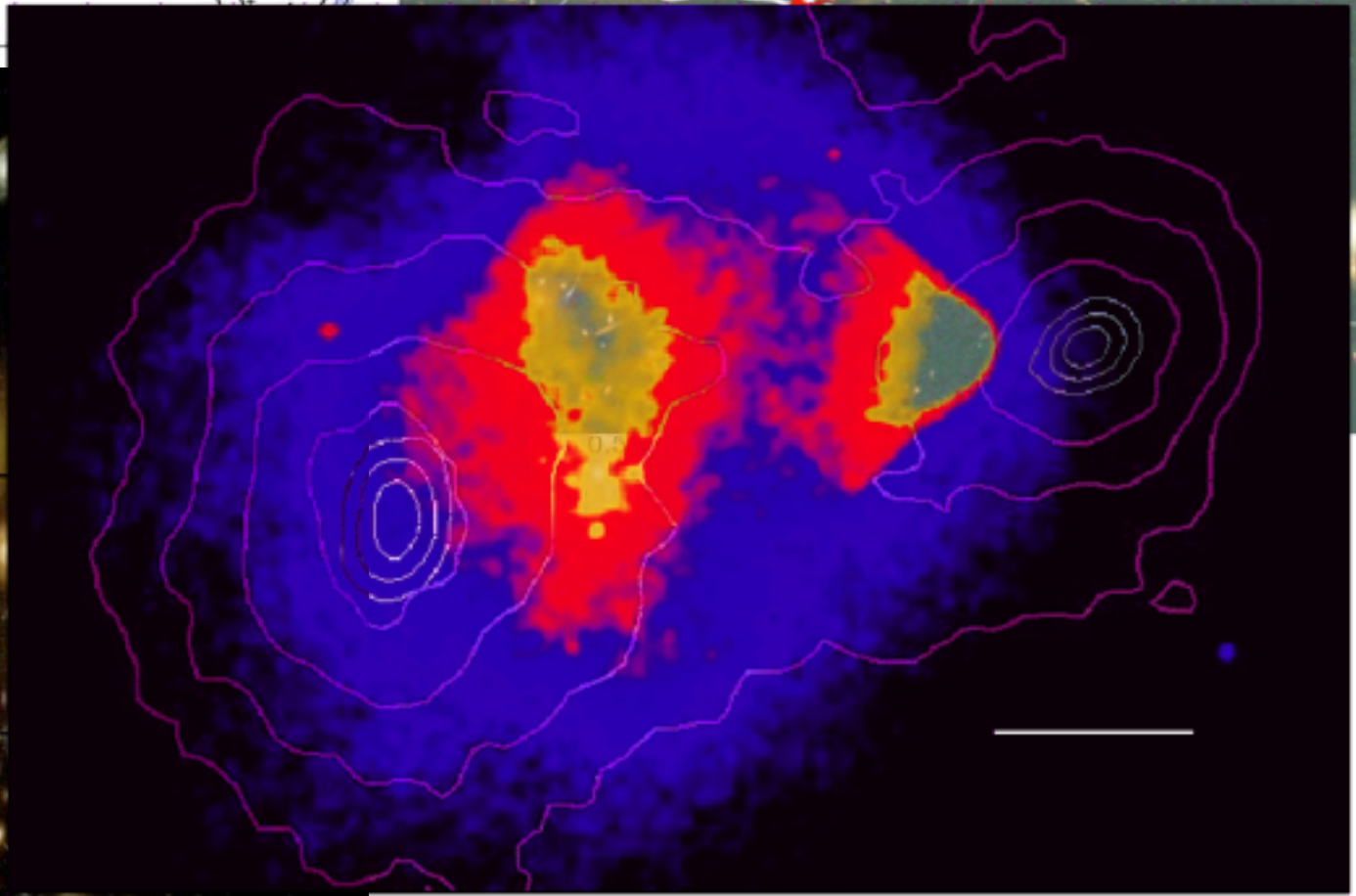
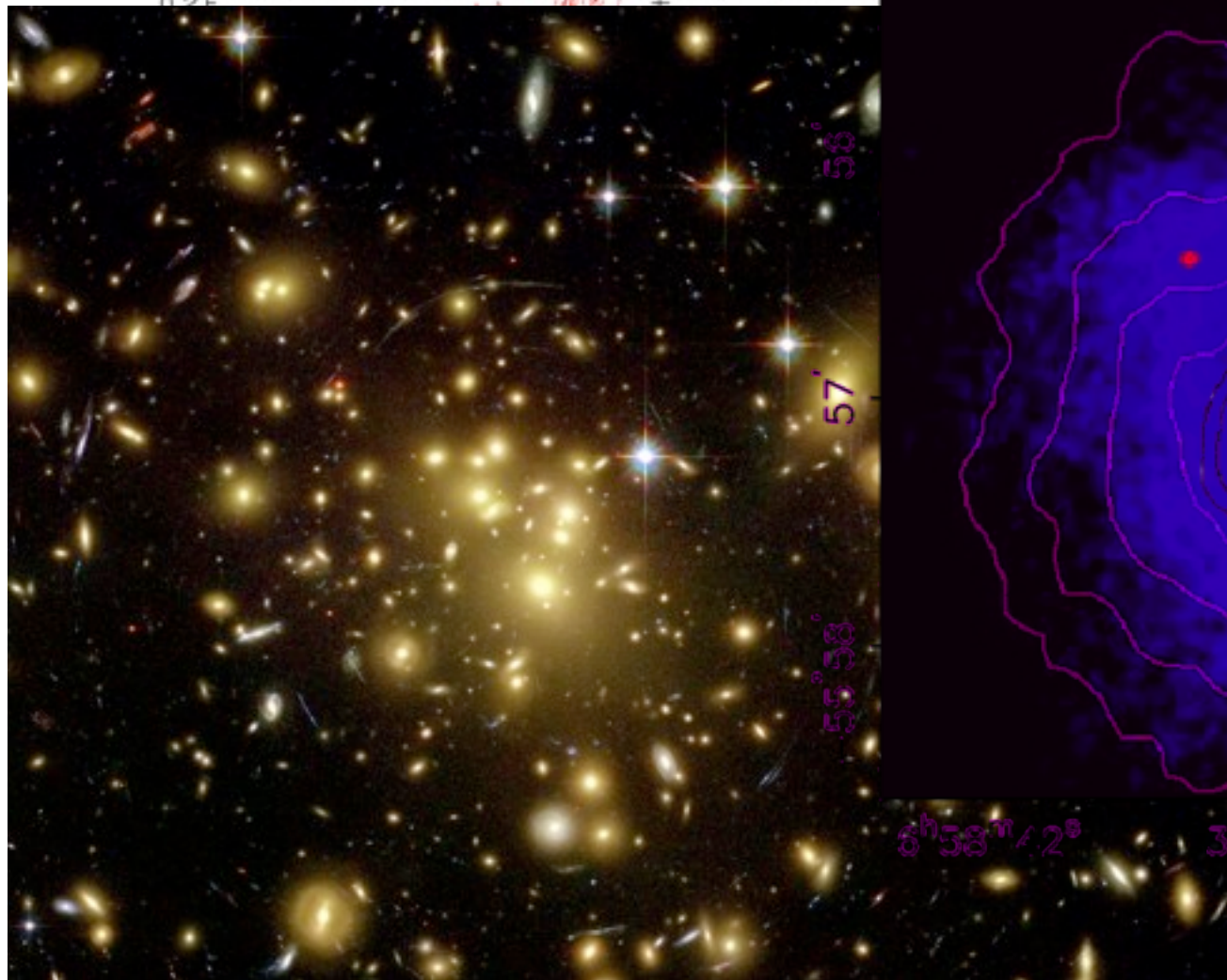
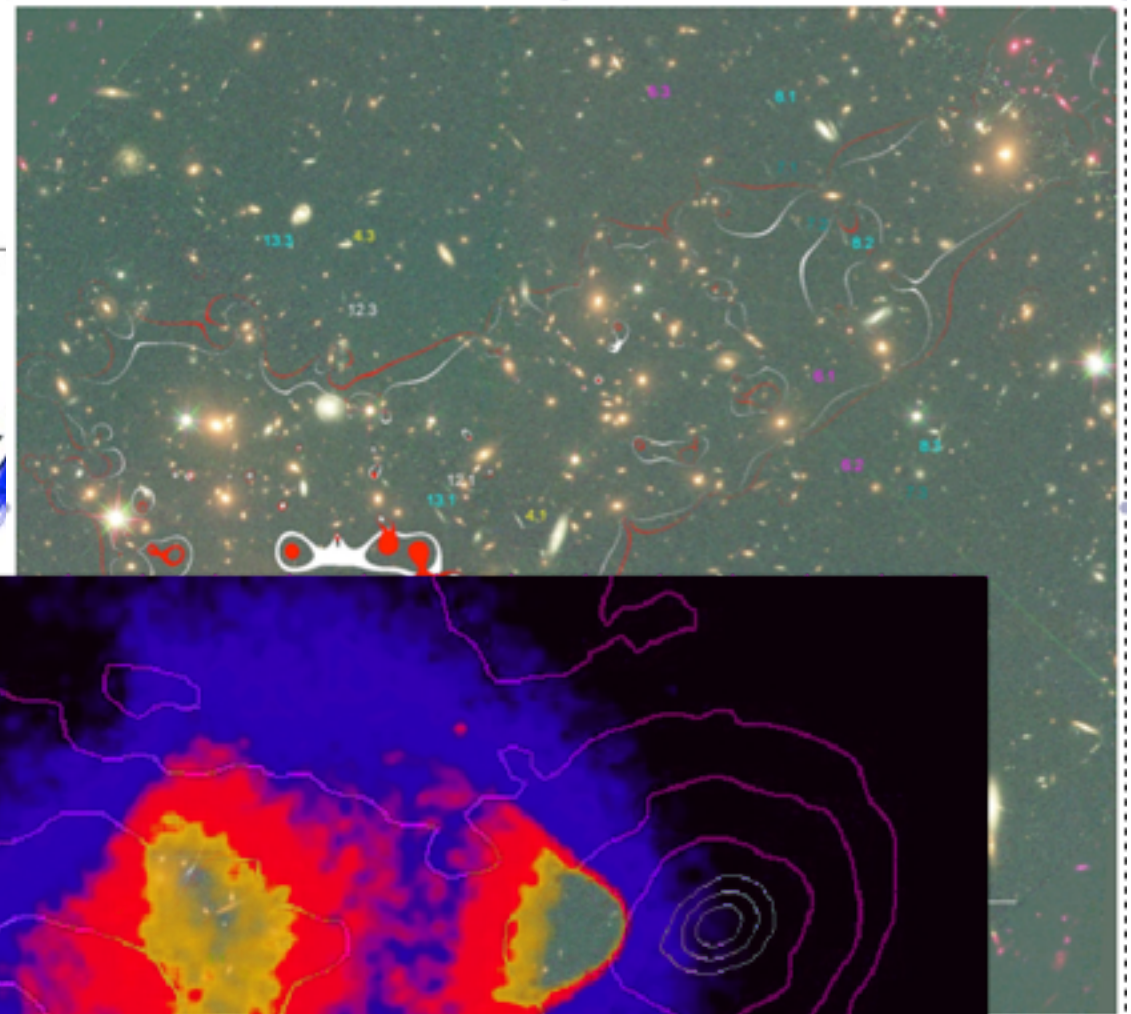
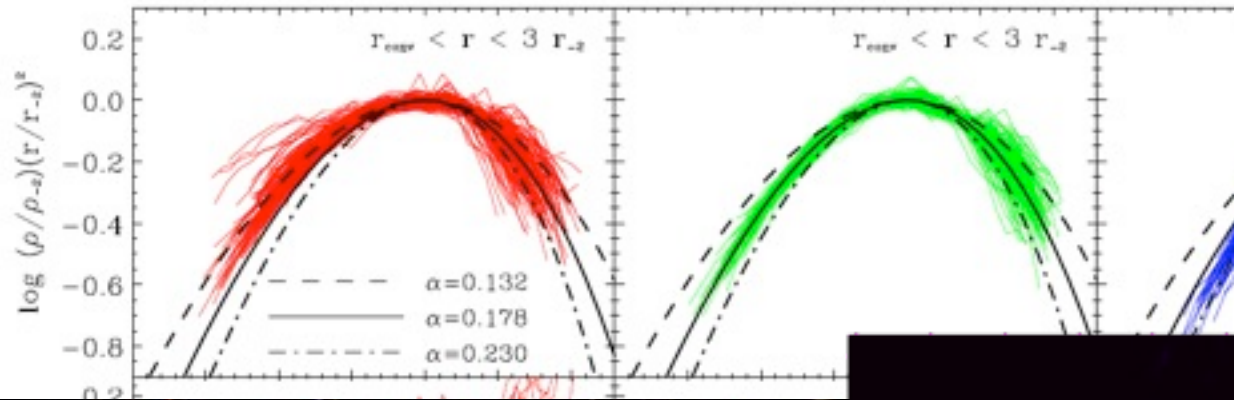
e.g. Broadhurst08+

Zitrin10+

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Cluster puzzles



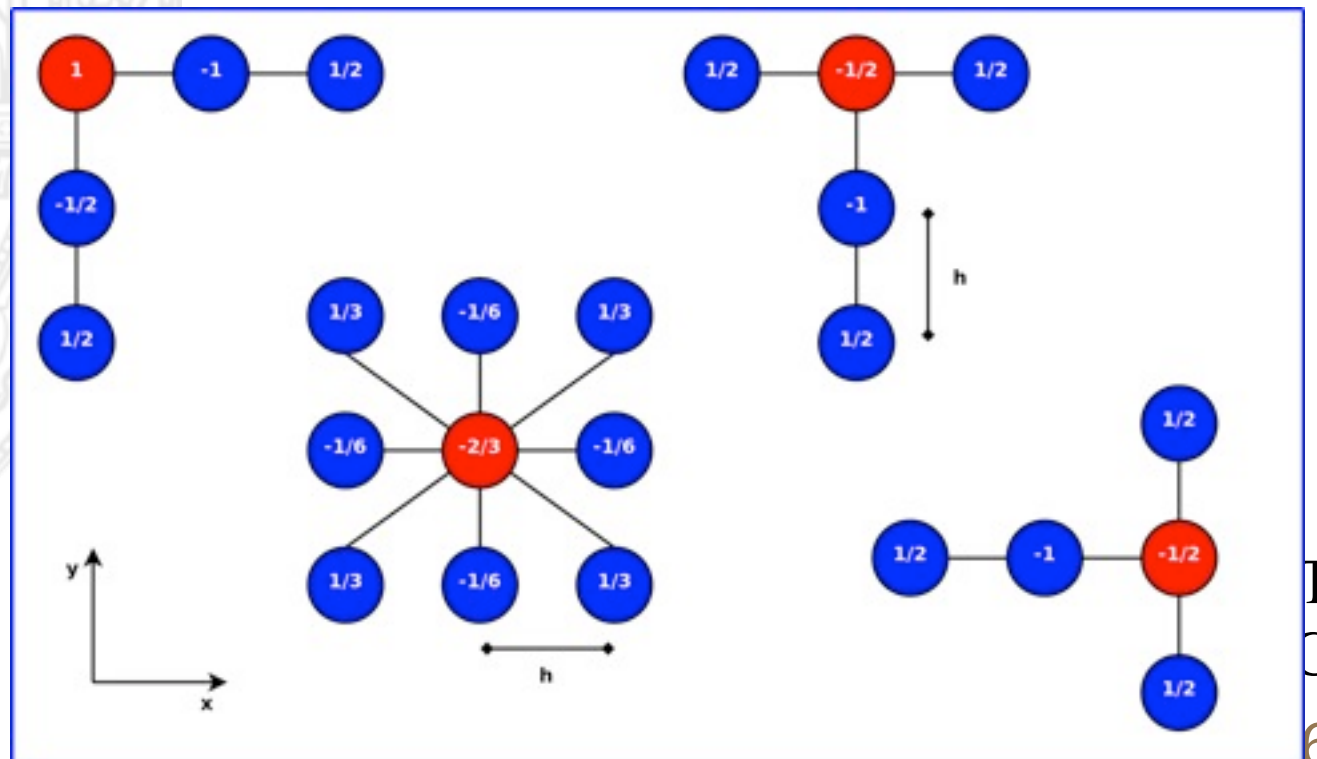
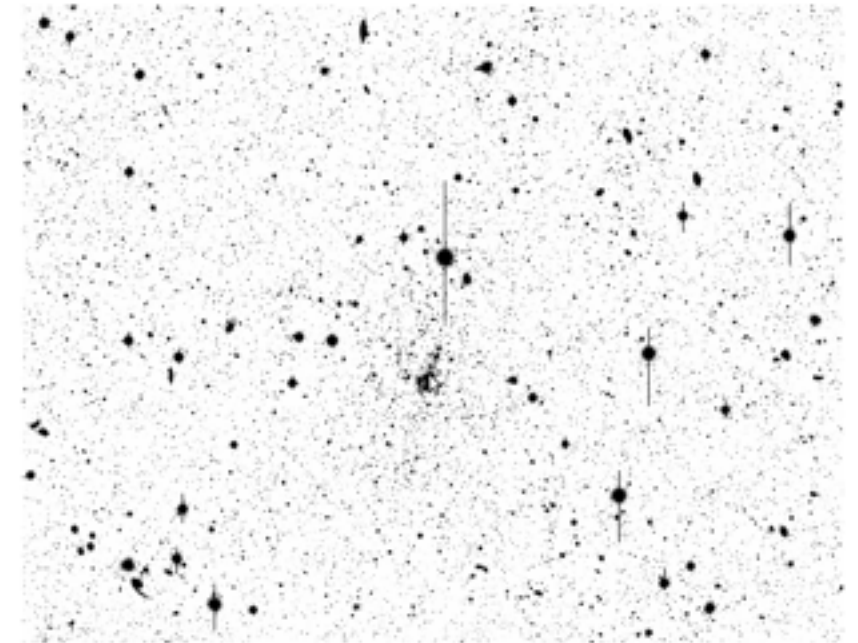
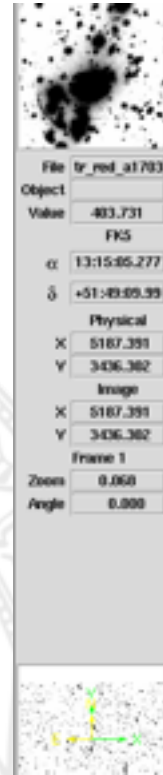
e.g. Broadhurst08+

e.g. Lee10+

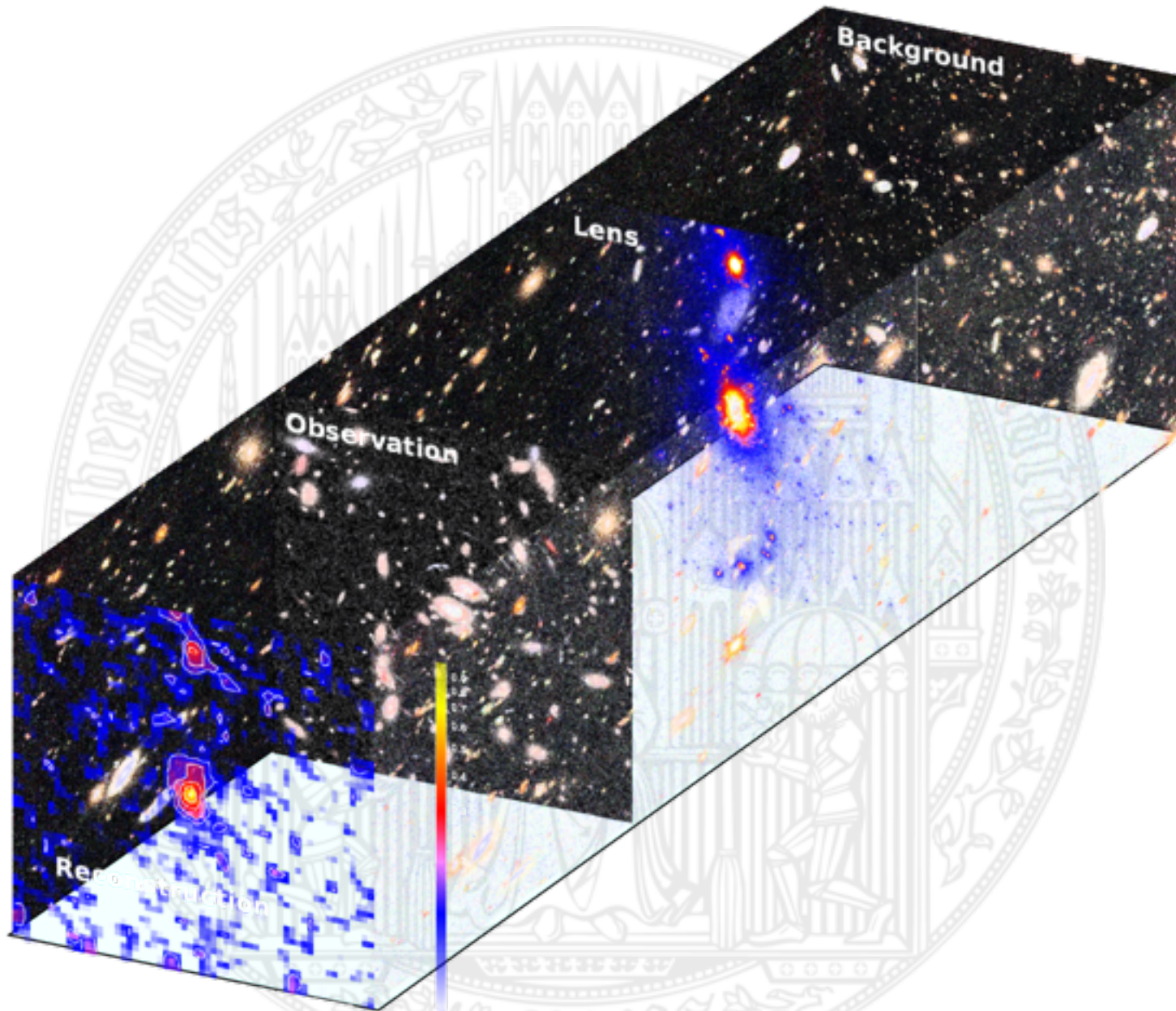
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Joint lensing reconstructions

- ▶ Fully nonparametric approach
- ▶ $\chi^2(\psi) = \chi_w^2(\psi) + \chi_s^2(\psi) + \dots$
- ▶ Adaptive-grid method
- ▶ Does not assume light traces mass
- ▶ Easily extendable
- ▶ Future: Kinematics, gas constraints



Simulating lensing

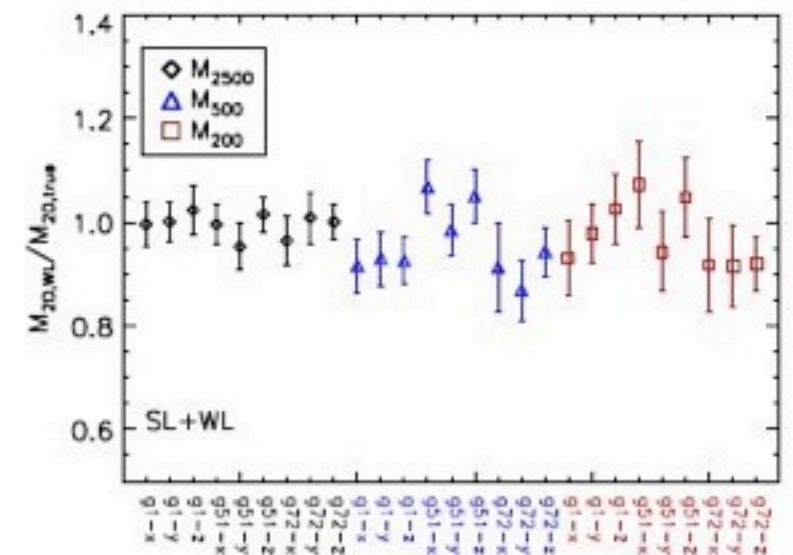
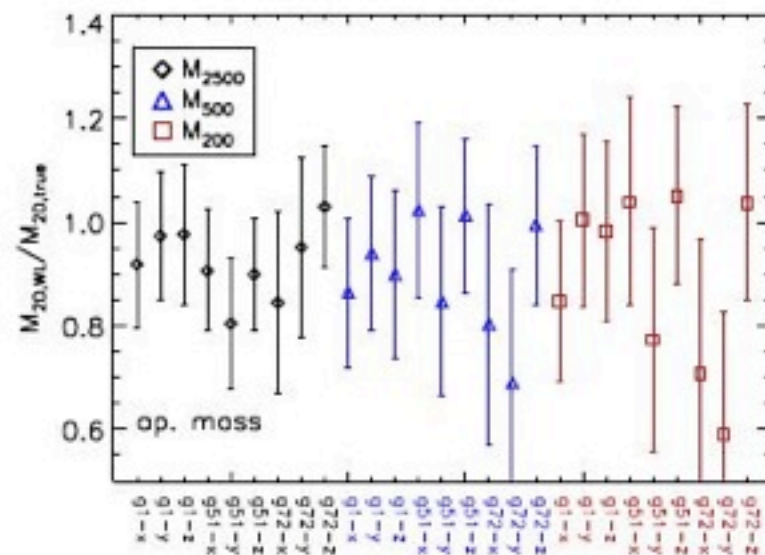
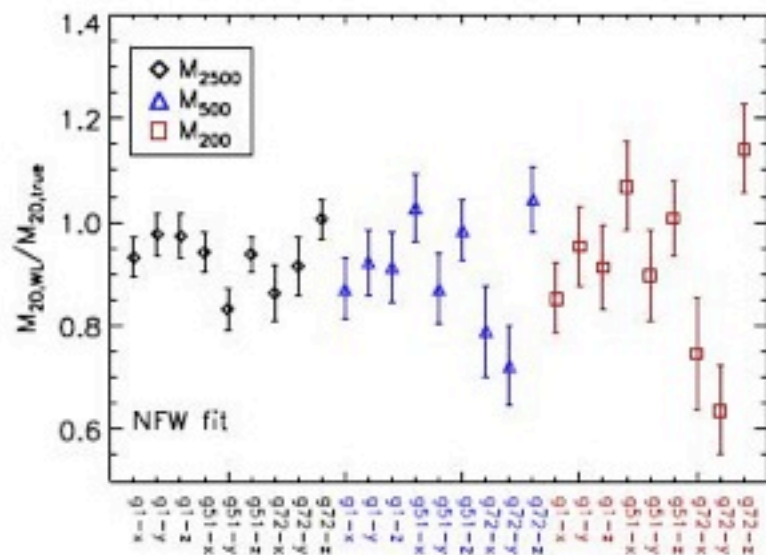


Name	Description
D	aperture diameter
g	detector gain
A_{pix}	pixel area
$F(\lambda)$	used filter
$M(\lambda)$	mirror filter curve
$O(\lambda)$	optics filter curve
$C(\lambda)$	CCD filter curve
FoV	total field-of-view
RON	detector readout-noise
f	flat-field accuracy
a	residual flat-field error
PSF	PSF model
t_{exp}	exposure time
$A(\lambda)$	atmospheric extinction
m_a	airmass
SED_{sky}	sky-background emission
SED_{gal}	background population
α	deflection angle map

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Simulating lensing

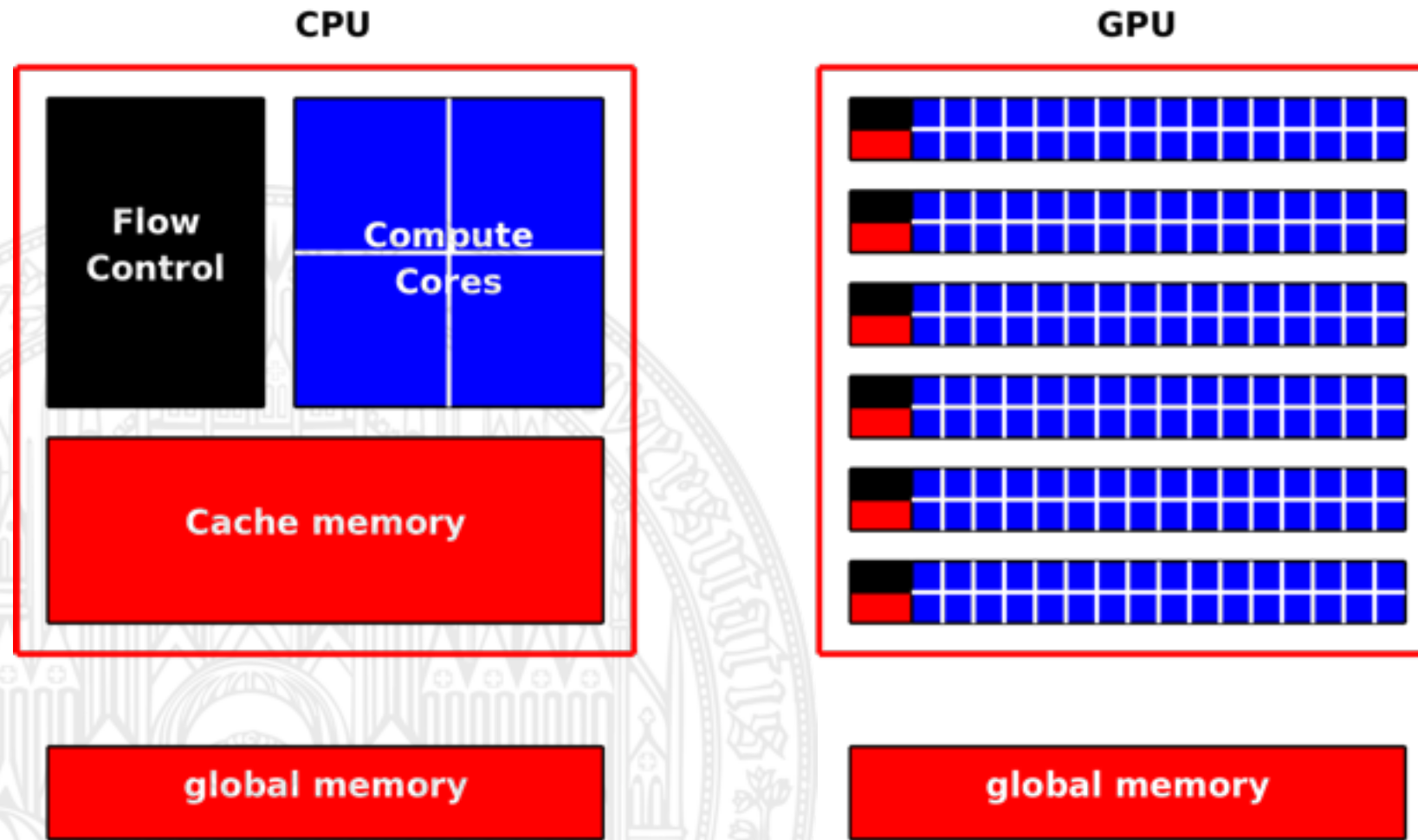


Meneghetti, Rasia, JM et al. 2010

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Modern implementation: GPUs



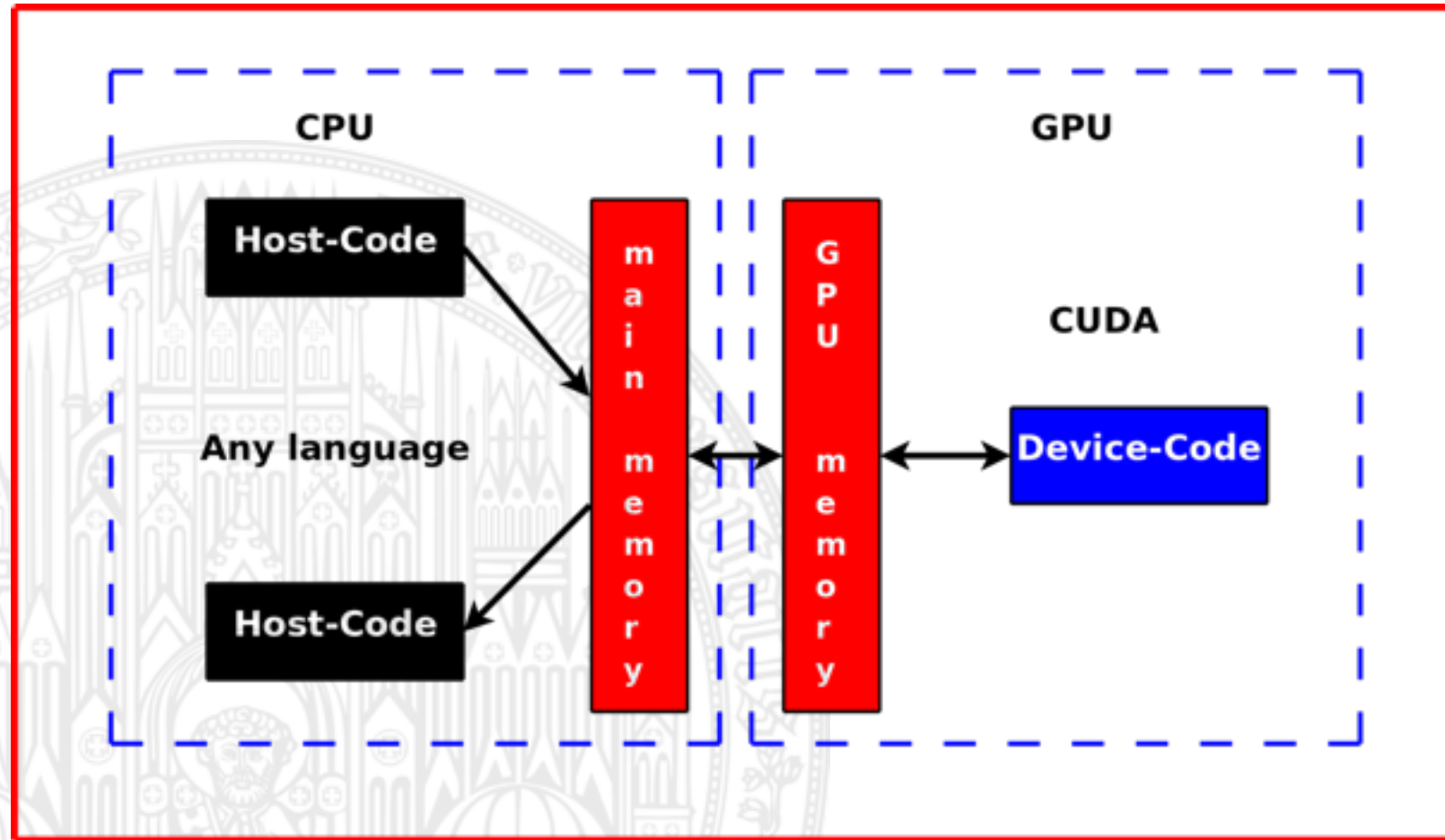
- ▶ Respect SIMT: hundreds/thousands of threads
- ▶ Respect memory hierarchy
- ▶ Almost no need for synchronisation
- ▶ Balance of parallelisation and work per thread

- ▶ Use CUDA C or OpenCL (Some alternatives upcoming)
- ▶ Available libraries: CUBLAS, CUFFT, CUSPARSE, CURAND, CULA
- ▶ Combination with MPI possible
- ▶ Some child diseases

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GPU systems and speed-up

NVIDIA TESLA C2050 (Fermi)

1 TFLOP peak (single)

500 GFLOP (double)

3 GB ECC global memory

14 (16) Multi-processors

448 (512) streaming cores



Our current workhorse: Jabba

2x quadcore XEON @ 2.4 Ghz

16 GB ECC main memory

NVIDIA TESLA C1060

NVIDIA TESLA C2050

Cluster performance for **5000** EUR

Speed-up

$$B_{lk} = a_i b_j C_{ij} D_{ik} E_{jl}$$

one core CPU: 82.3 s

240 core GPU: **1.03** s

+ ITA/ZITI cluster Kolob

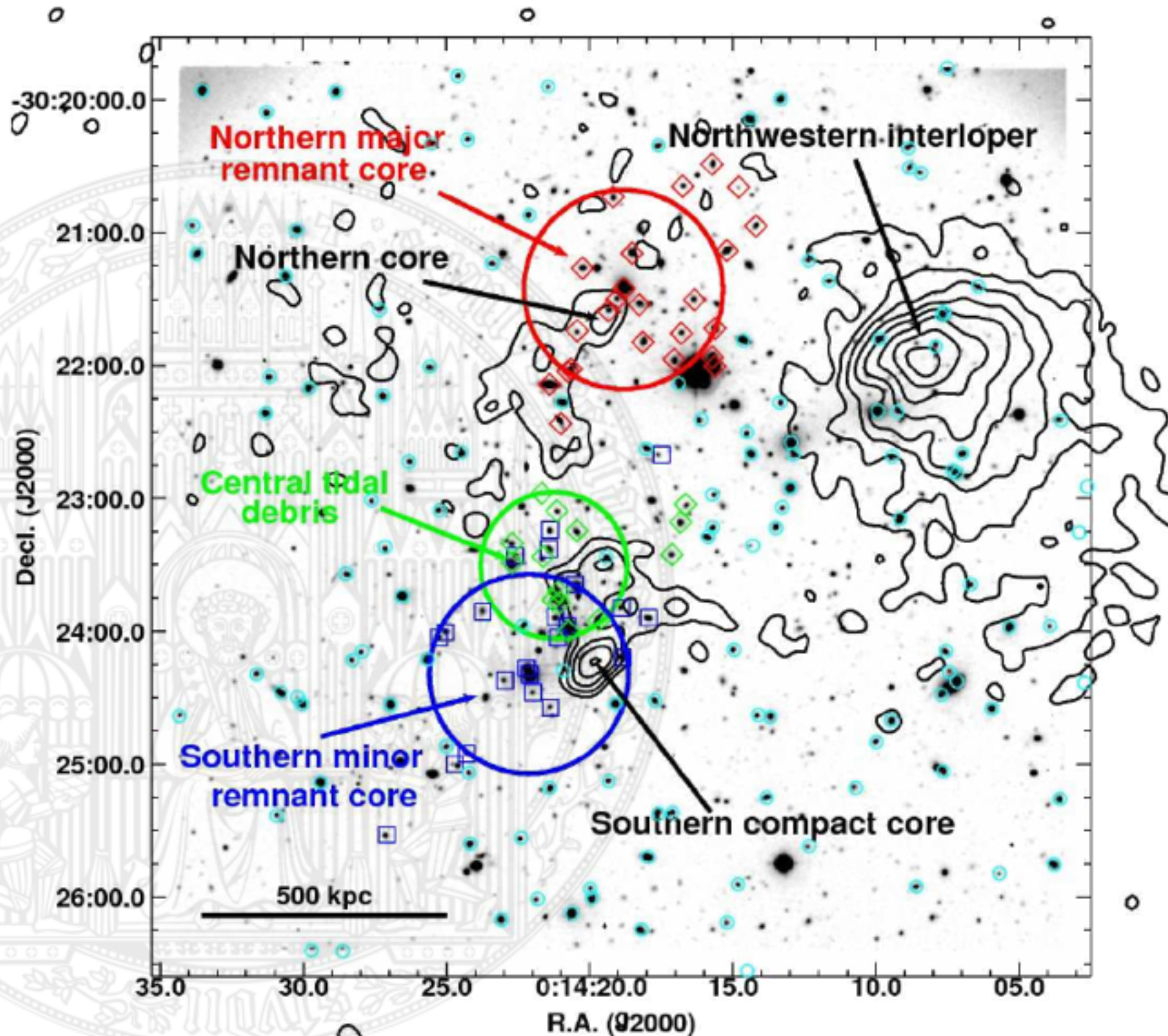
+ Granted SFB

machine

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One of the pearls: Abell 2744



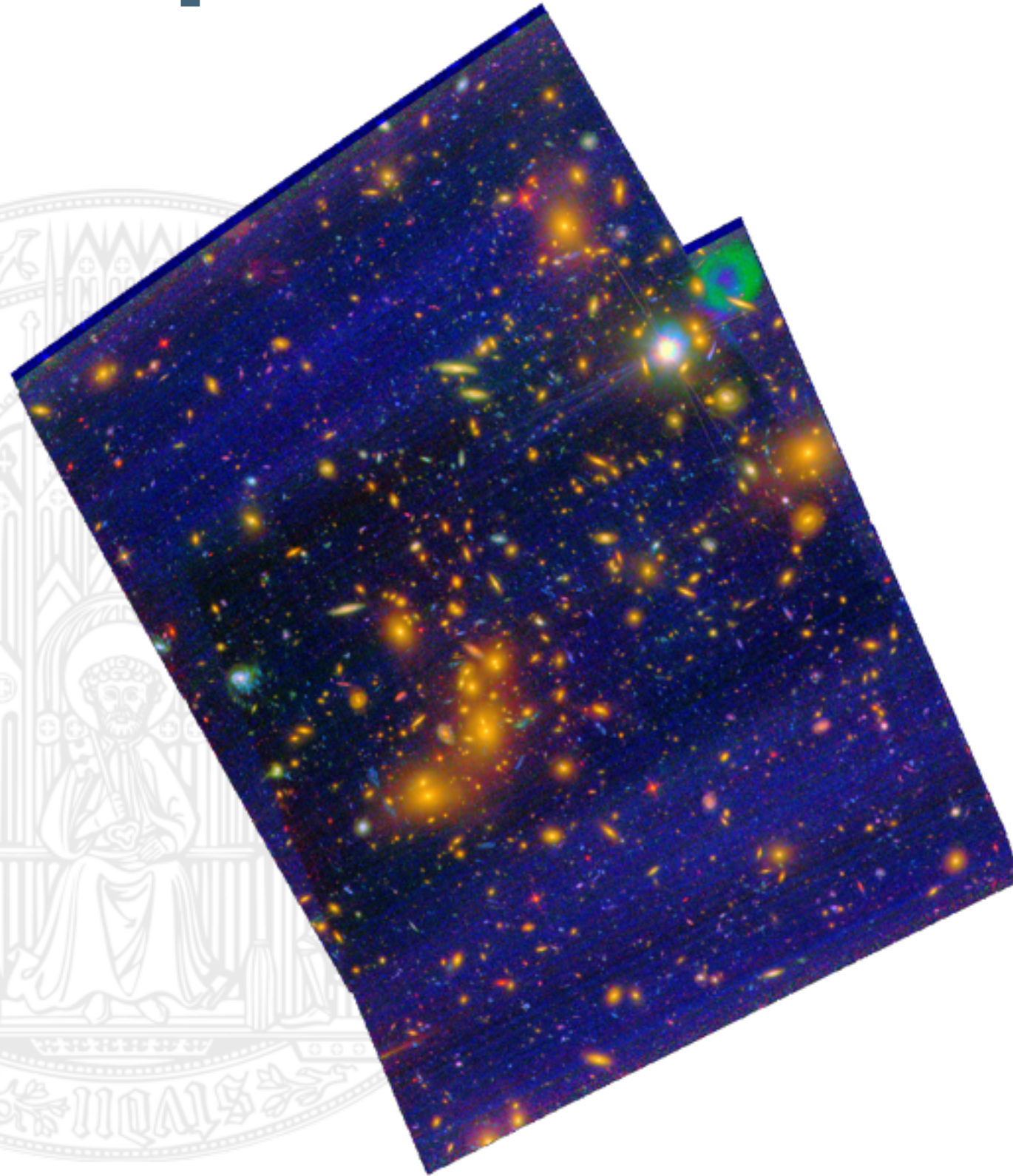
Owers10+

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One of the pearls: Abell 2744

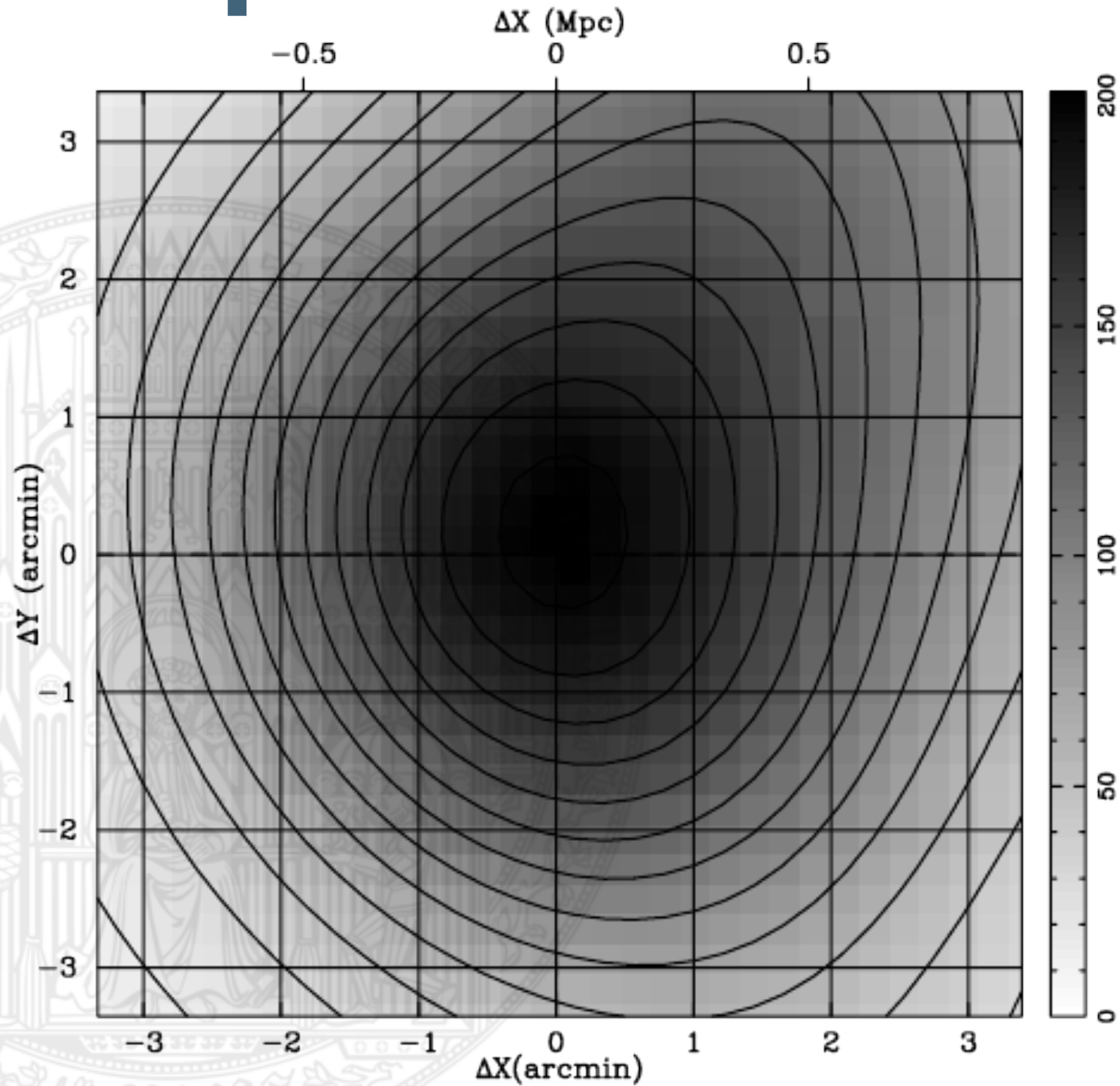


3band HST/ACS imaging, PI: R. Dupke

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One of the pearls: Abell 2744

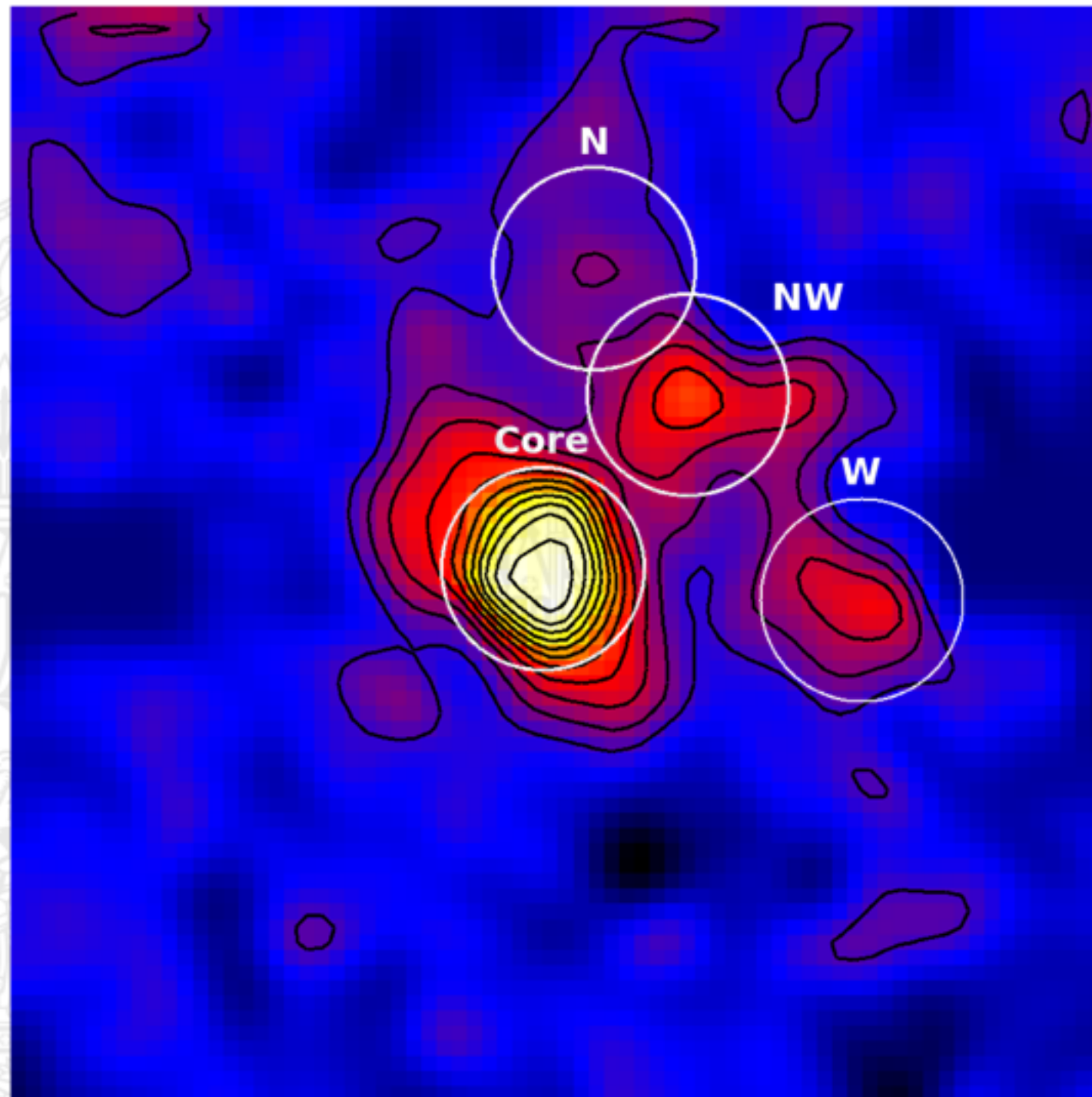


Cypriano 04+

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One of the pearls: Abell 2744

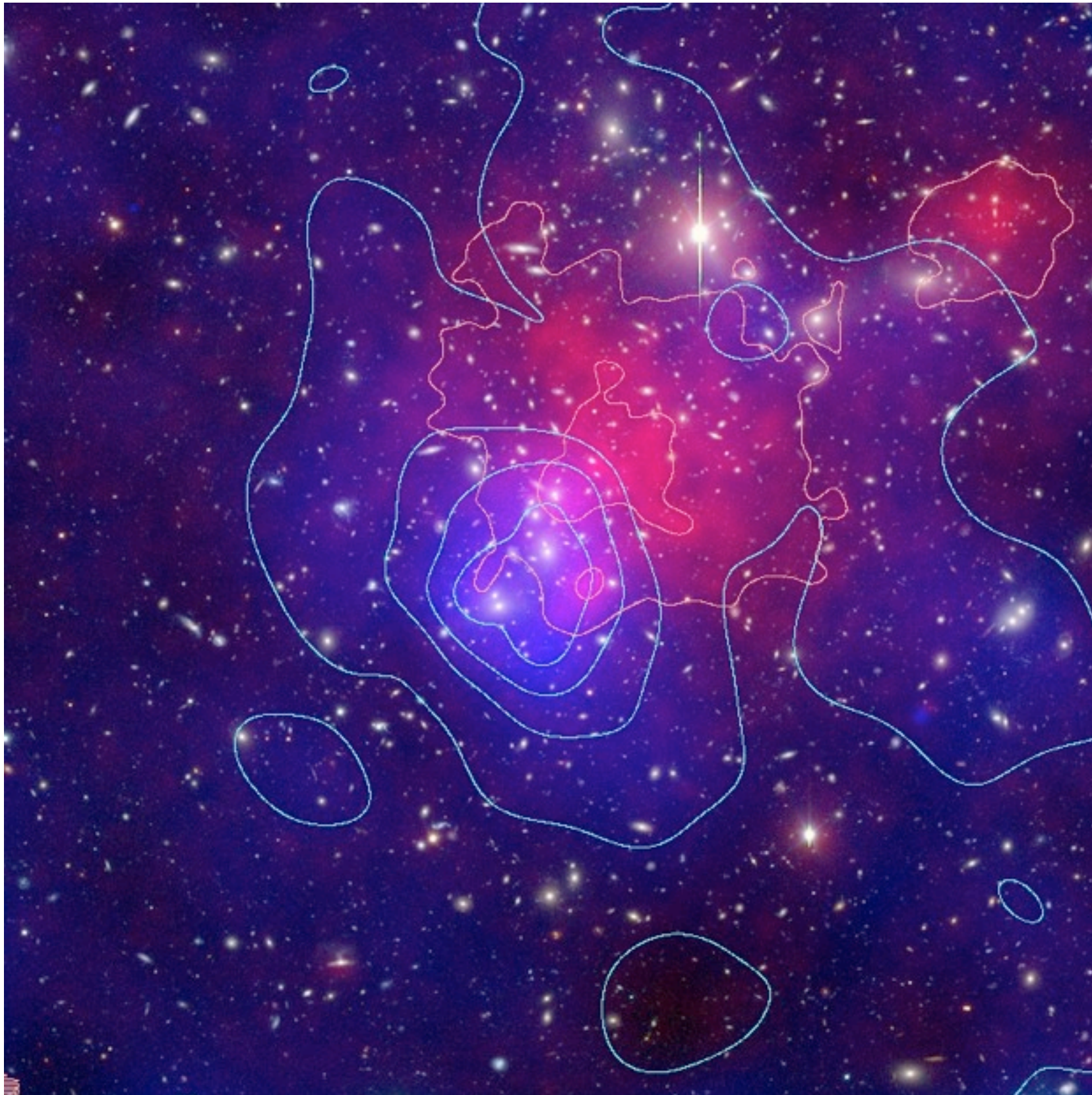


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JM11+ in prep.

Creation of cosmic structure



**JM11+ in prep.
(next week)**

**D. Coe (STScI)
R. Dupke (Michigan)
R. Massey (Edinburgh)
A. Zitrin (Tel Aviv)
T. Broadhurst (Bilbao)
E. Cypriano (S. Paolo)
N. Okabe (Taipei)
M. Meneghetti (Bologna)
J. Rhodes (JPL)
L. Moustakas (JPL)**

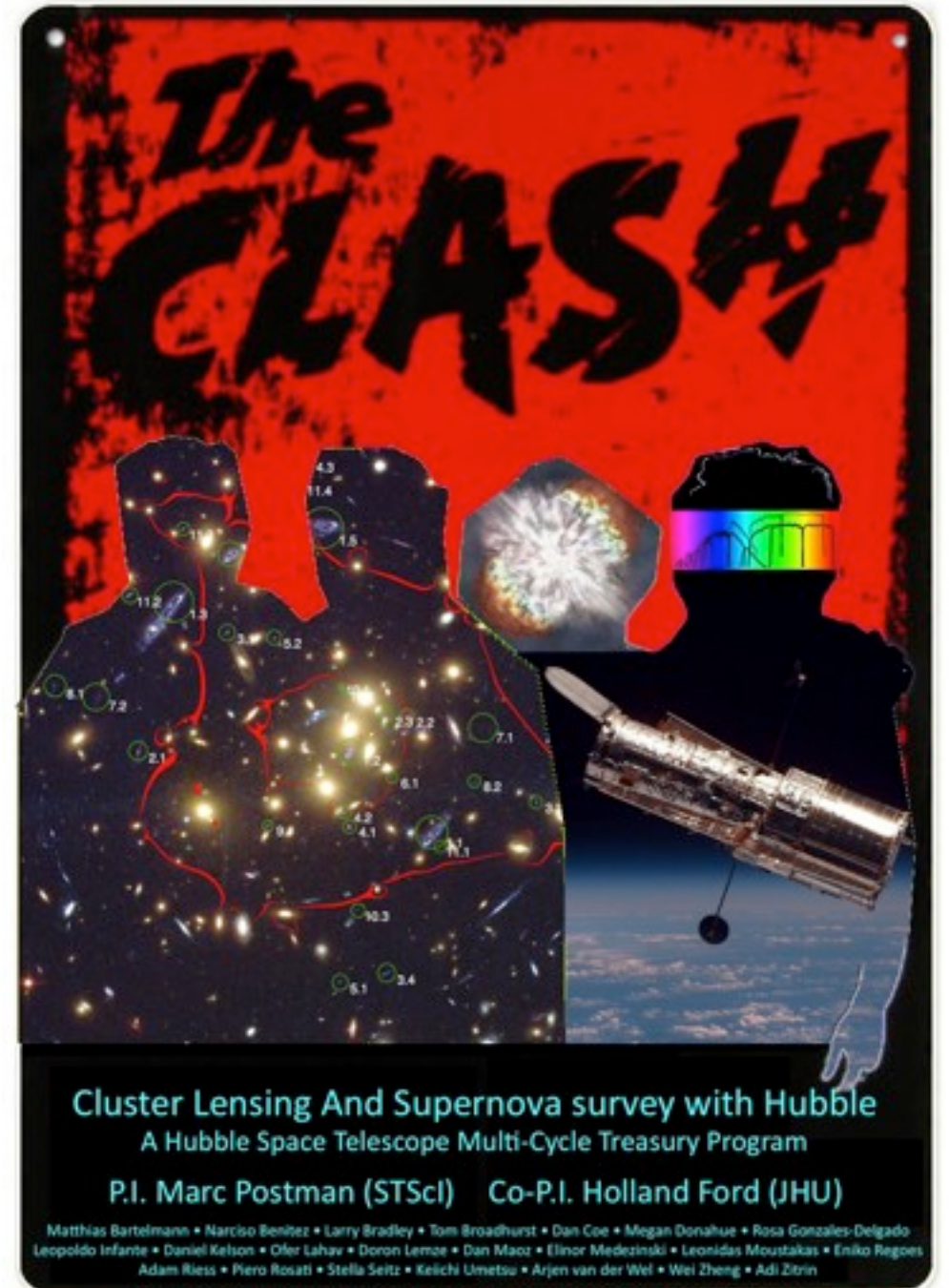
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CLASH: A HST/MCT programme

Science Drivers

- ▶ To map dark matter in galaxy clusters
- ▶ To detect SN out to redshifts $z > 1.5$
- ▶ To detect and characterise $z > 7$ galaxies
- ▶ To study the galaxies in and behind the clusters



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CLASH: A HST/MCT programme



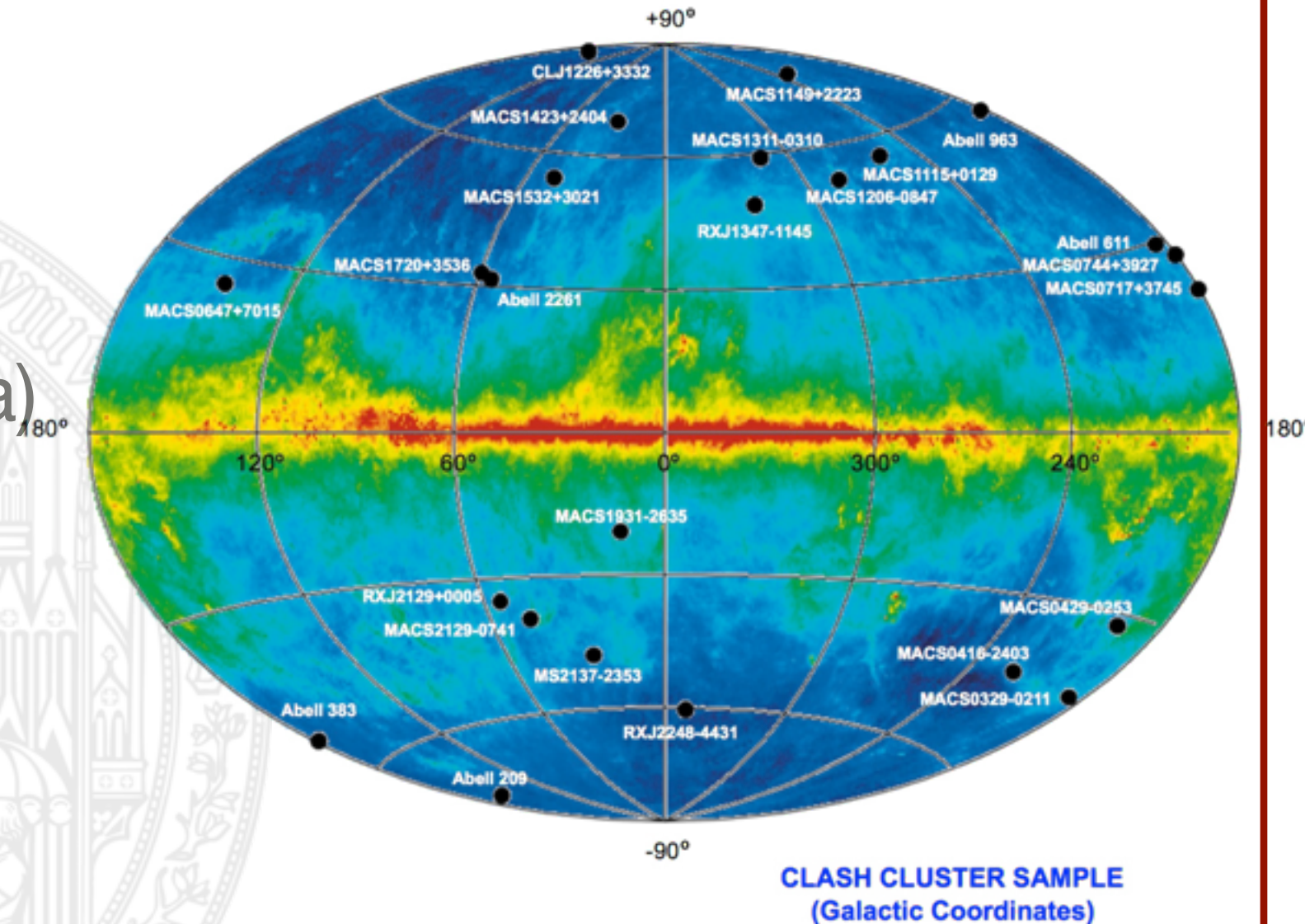
Granada, September 20, 2010

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TAT
ERG

The observations: **524** HST orbits

- ▶ 25 clusters
($0.18 < z < 0.9$)
- ▶ X-ray selected
(Chandra archival data)
- ▶ relaxed
- ▶ ACS/WFC3 in parallel
- ▶ 16 bands & 20 orbits
per cluster (NIR - NUV)

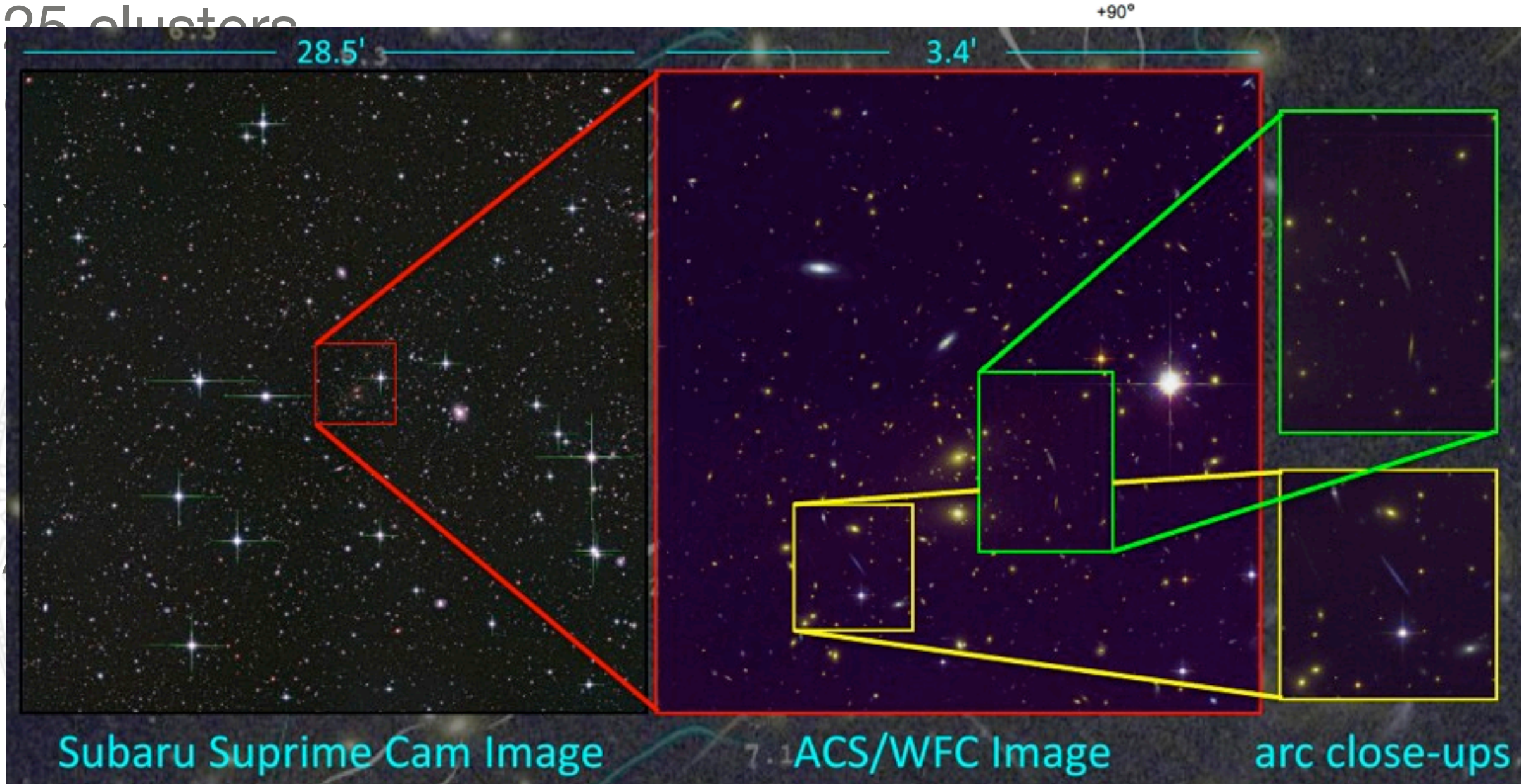


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The observations: 524 HST orbits

▶ 25 clusters



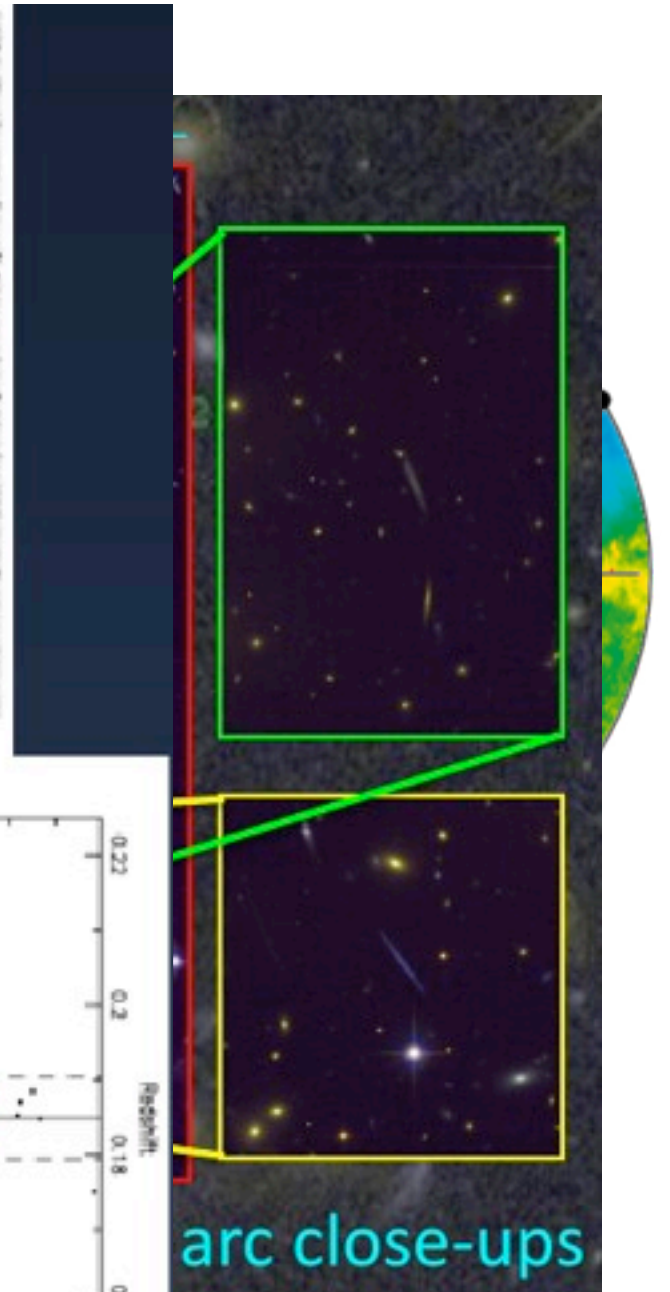
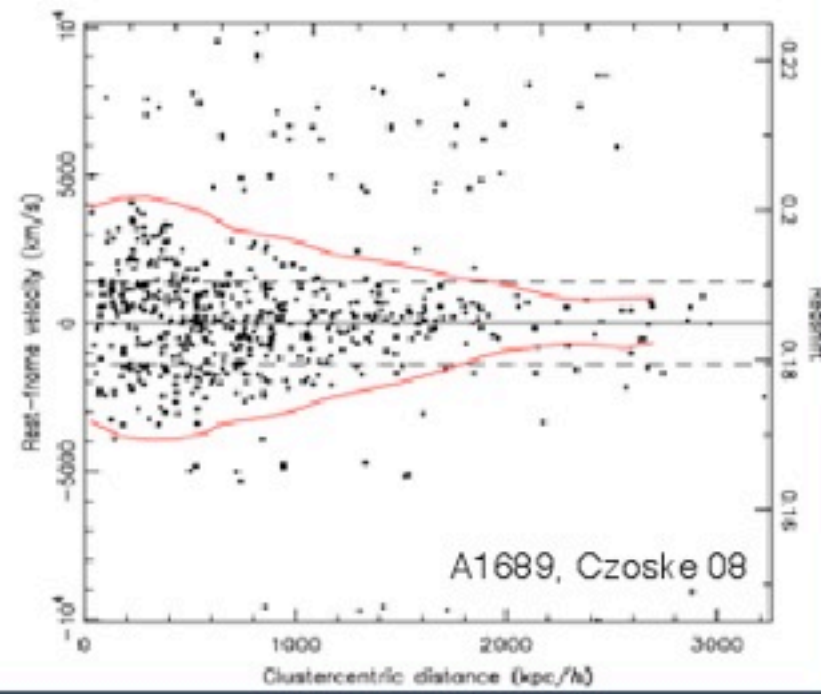
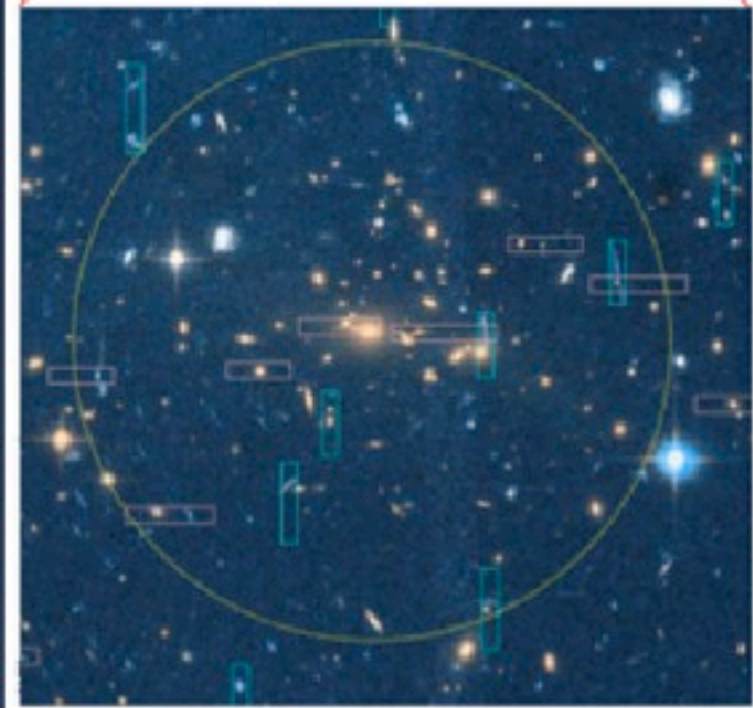
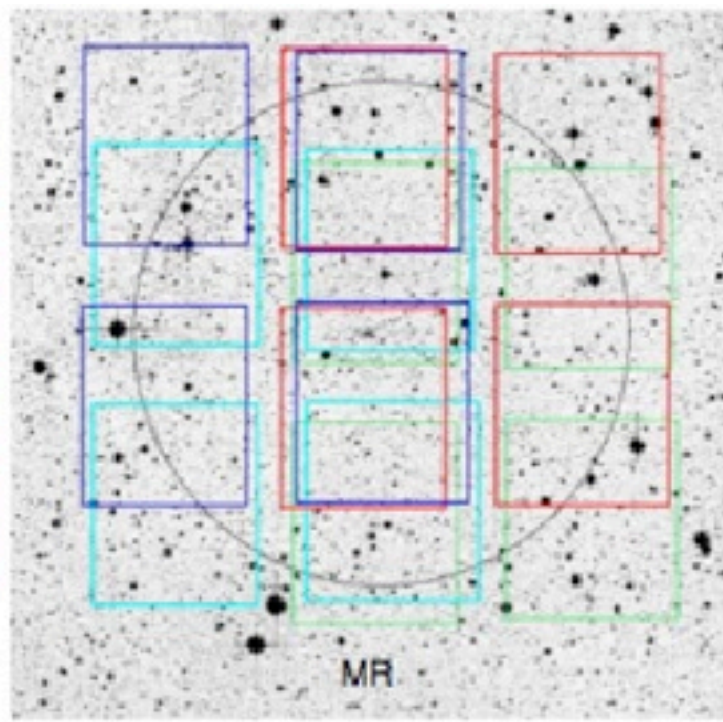
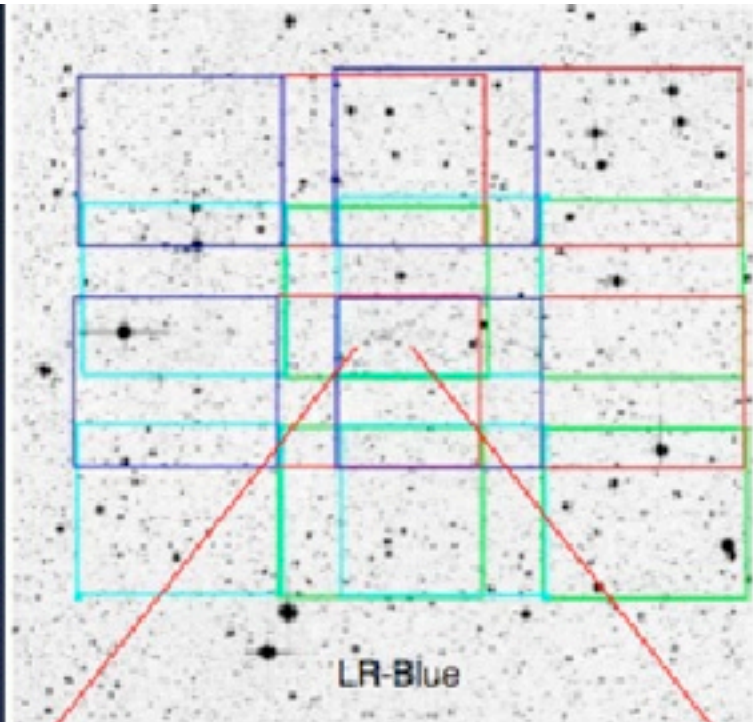
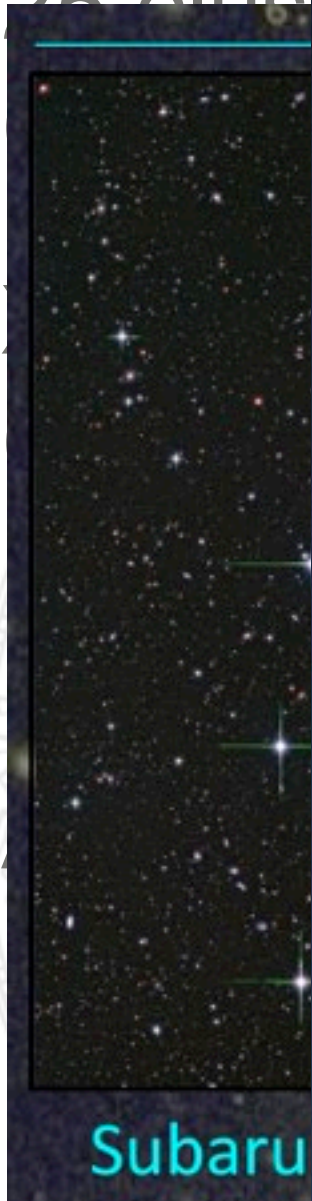
SUBARU BVRIZ weak lensing

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The observations: 524 HST orbits

▶ 25 plus

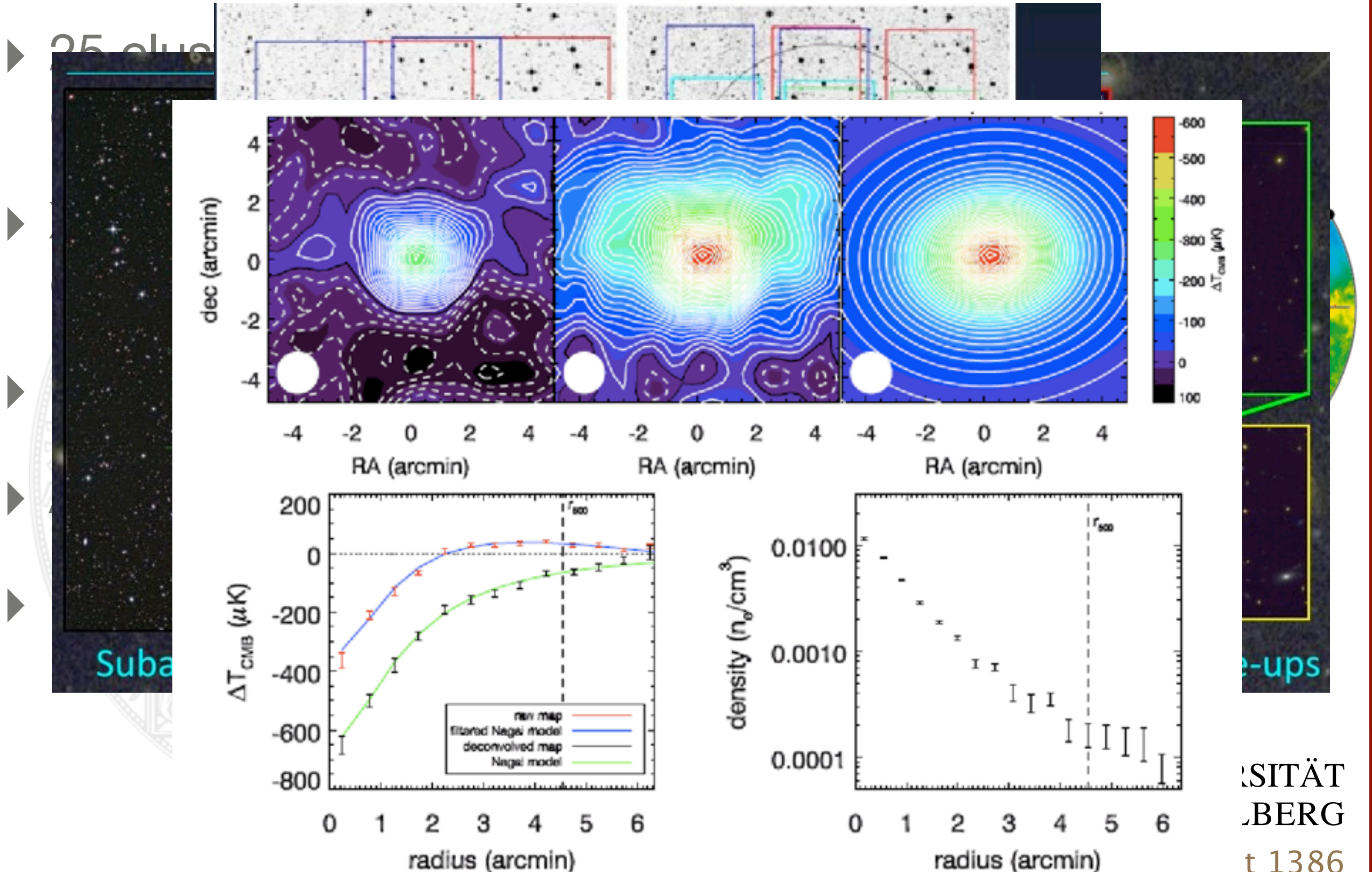


GTC/VLT/Magellan spectroscopy

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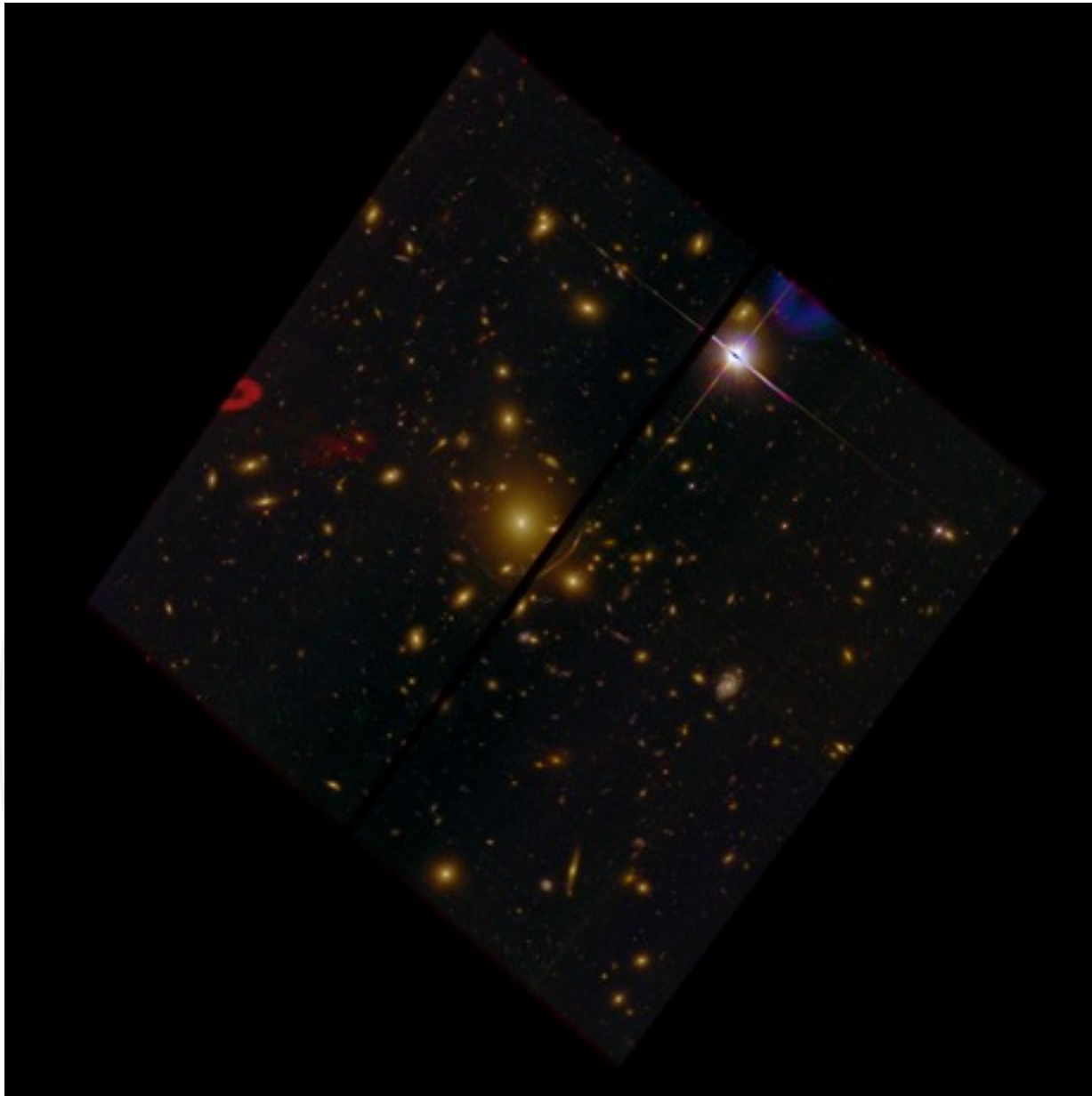
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The observations: 524 HST orbits

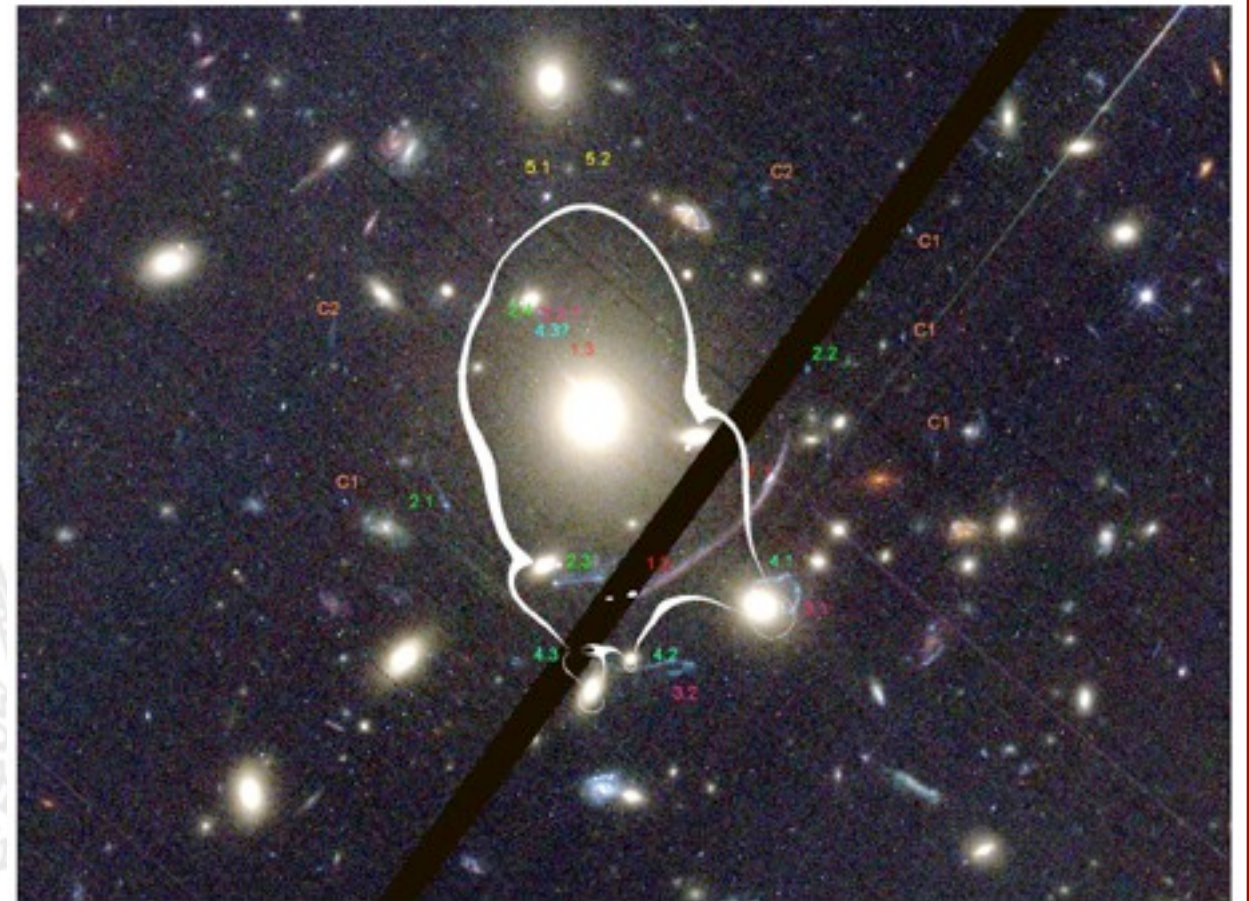


Bolocam/AMiBA/SZA/Mustang SZE observations

First results: Abell 383



**Colour composite using
Trilogity, credit Dan Coe**

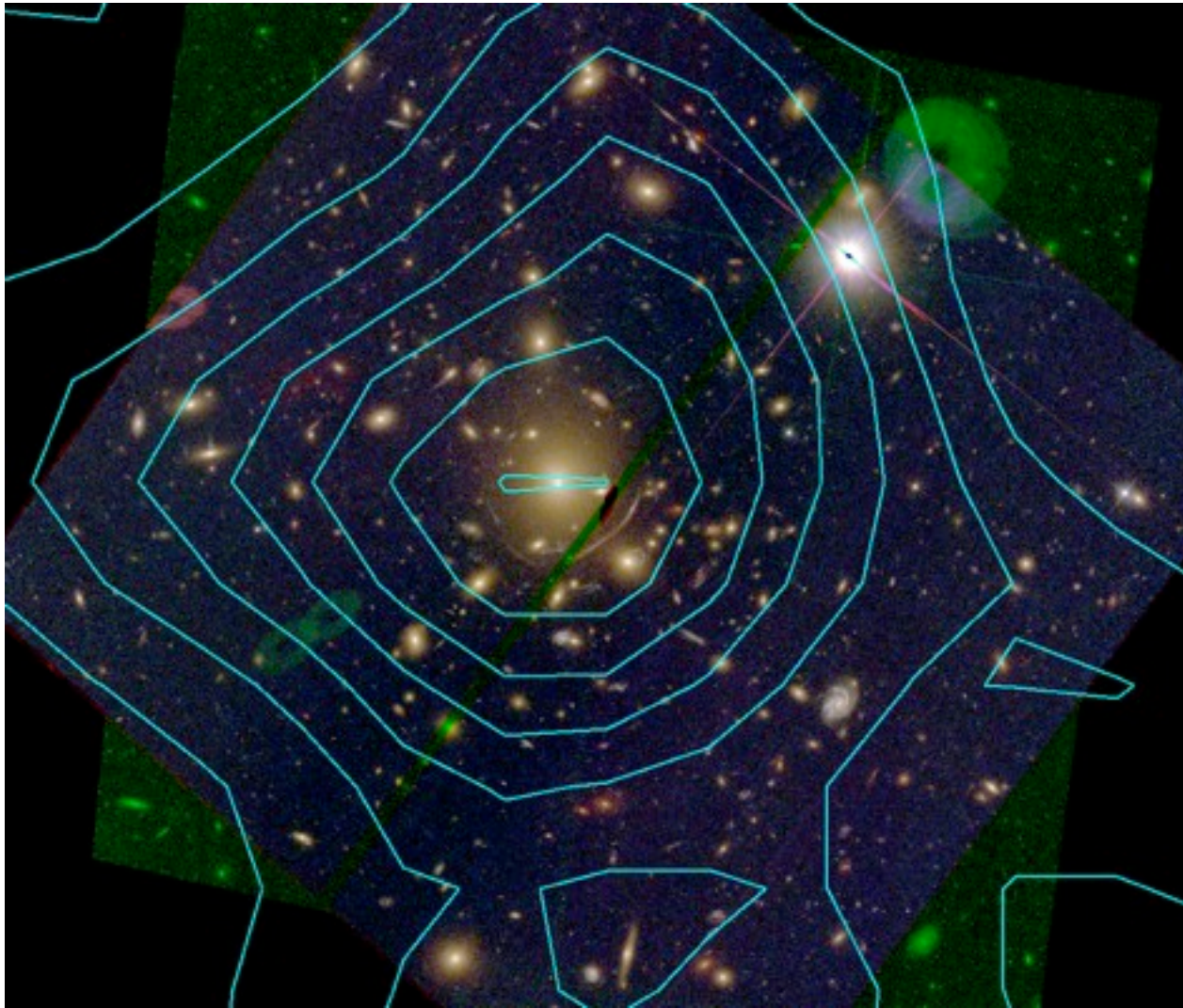


**First SL mass models,
Zitrin11+ in prep.**

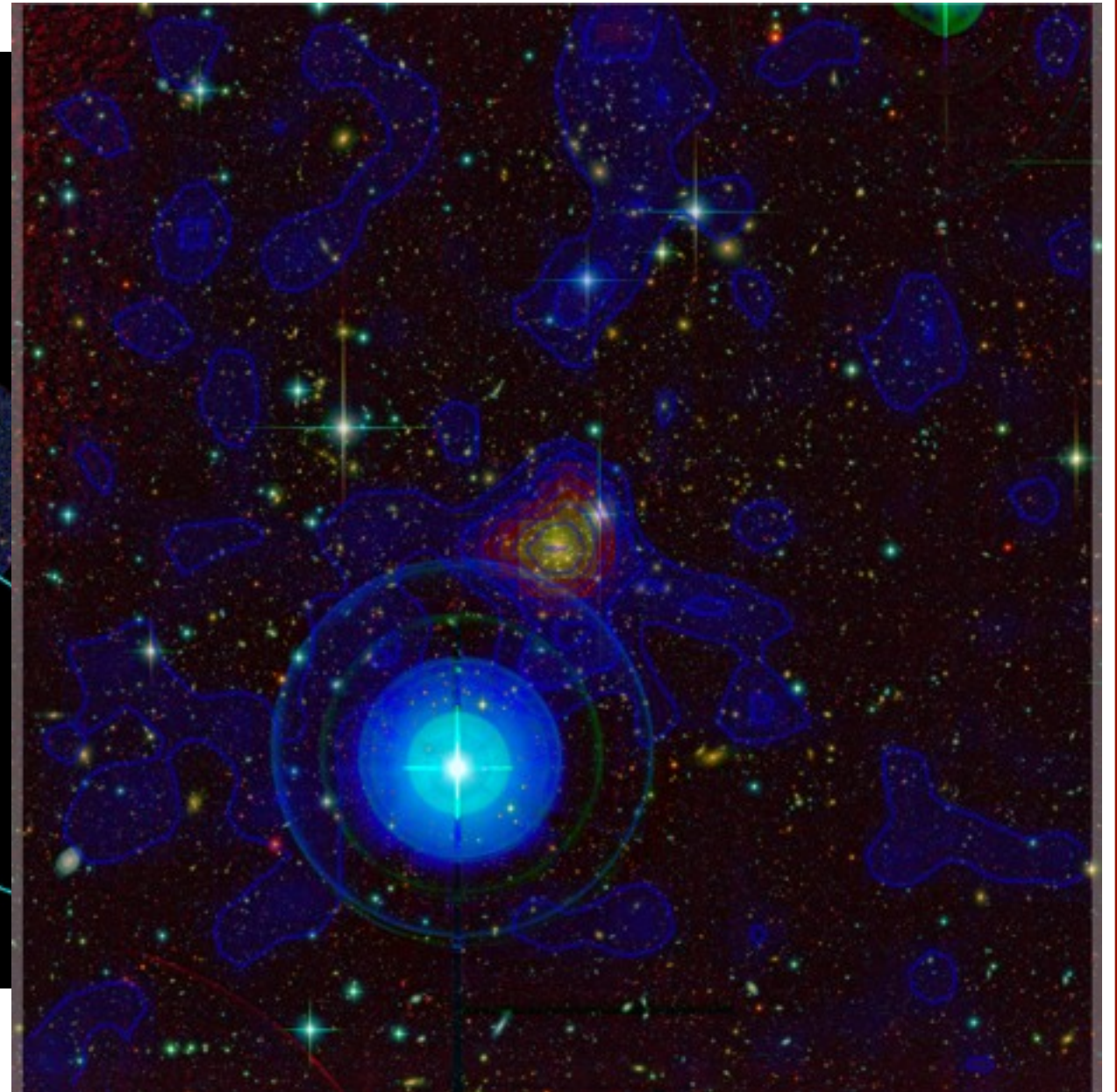
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First results: Abell 383



WL contours on ACS (JM)



WL contours on SUBARU (JM)

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Conclusions

- ▶ Galaxy clusters are an exciting topic for cosmology
- ▶ Combined lensing codes are able to reconstruct their mass distribution to a few percent
- ▶ For testing and comparing to simulations, the full pipeline has to be simulated
- ▶ GPUs are an energy -and price-efficient way to implement time-consuming codes
- ▶ Abell 2744: cosmic structure formation at work
- ▶ CLASH will help answering some open questions, at least we hope so

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